



Geo-neutrino Measurement with KamLAND

Hiroko Watanabe

Research Center for Neutrino Science (Tohoku Univ.)
for the KamLAND Collaboration

Contents

1. Introduction
2. Geo-neutrino Measurement Results
3. Future Prospects
4. Summary

Contents

1. Introduction
2. Geo-neutrino Measurement Results
3. Future Prospects
4. Summary

A. Gando,¹ Y. Gando,¹ H. Hanakago,¹ H. Ikeda,¹ K. Inoue,^{1,2} K. Ishidoshiro,¹ H. Ishikawa,¹ M. Koga,^{1,2}
R. Matsuda,¹ S. Matsuda,¹ T. Mitsui,¹ D. Motoki,¹ K. Nakamura,^{1,2} A. Obata,¹ A. Oki,¹ Y. Oki,¹
M. Otani,¹ I. Shimizu,¹ J. Shirai,¹ A. Suzuki,¹ Y. Takemoto,¹ K. Tamae,¹ K. Ueshima,¹ H. Watanabe,¹
B.D. Xu,¹ S. Yamada,¹ Y. Yamauchi,¹ H. Yoshida,¹ A. Kozlov,² S. Yoshida,³ A. Piepke,^{2,4} T.I. Banks,⁵
B.K. Fujikawa,^{2,5} K. Han,⁵ T. O'Donnell,⁵ B.E. Berger,⁶ J.G. Learned,⁷ S. Matsuno,⁷ M. Sakai,⁷ Y. Efremenko,^{2,8}
H.J. Karwowski,⁹ D.M. Markoff,⁹ W. Tornow,⁹ J.A. Detwiler,¹⁰ S. Enomoto,^{2,10} and M.P. Decowski^{2,5,11}

1 Research Center for Neutrino Science, Tohoku University

2 Kavli Institute for the Physics and Mathematics of the Universe (WPI),
University of Tokyo

3 Graduate School of Science, Osaka University

4 Department of Physics and Astronomy, University of Alabama

5 Physics Department, University of California, Berkeley, and Lawrence
Berkeley National Laboratory

6 Department of Physics, Colorado State University, Fort Collins

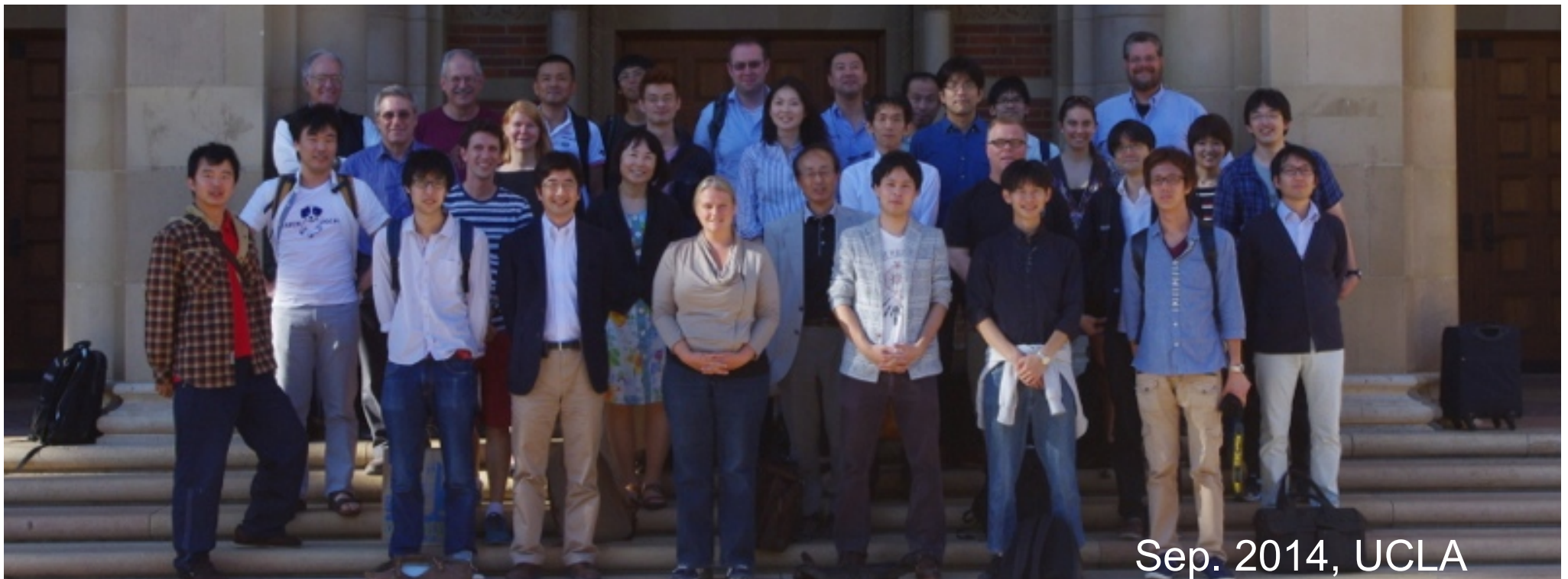
7 Department of Physics and Astronomy, University of Hawaii at Manoa

8 Department of Physics and Astronomy, University of Tennessee

9 Triangle Universities Nuclear Laboratory and Physics Departments at Duke
University and the University of North Carolina

10 Department of Physics, University of Wisconsin

11 Nikhef, Science Park 105, 1098 XG Amsterdam



Sep. 2014, UCLA

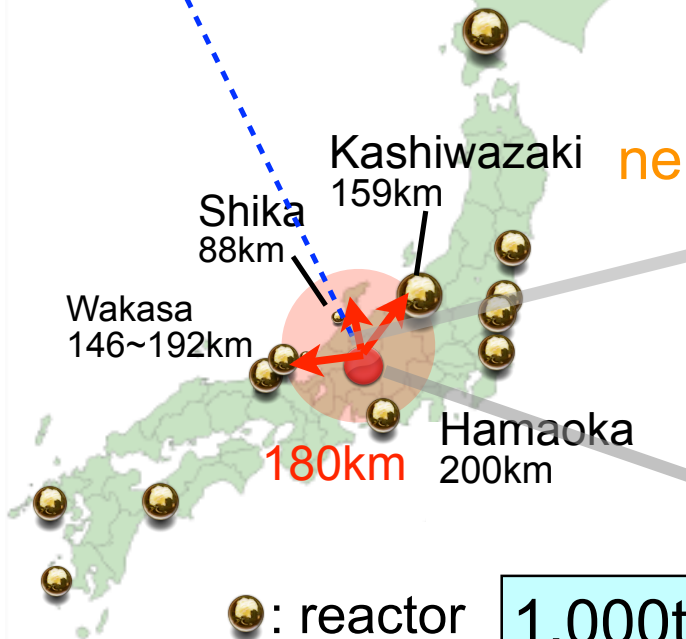
KamLAND

**Kamioka Liquid Scintillator
Anti-Neutrino Detector**

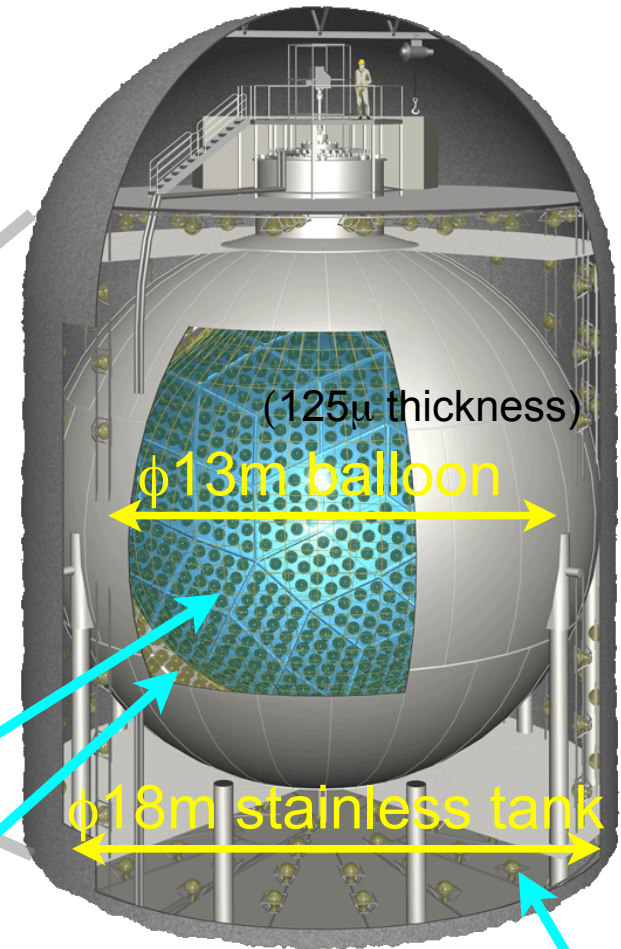
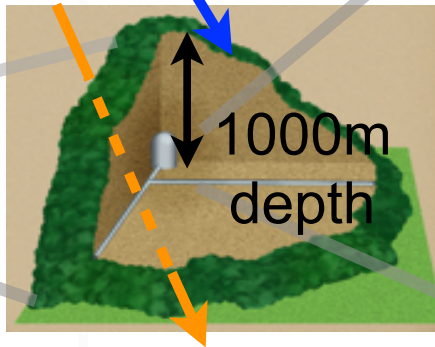
(operated since 2002)



Kamioka Mine



neutrino cosmic ray



1,000t Liquid Scintillator

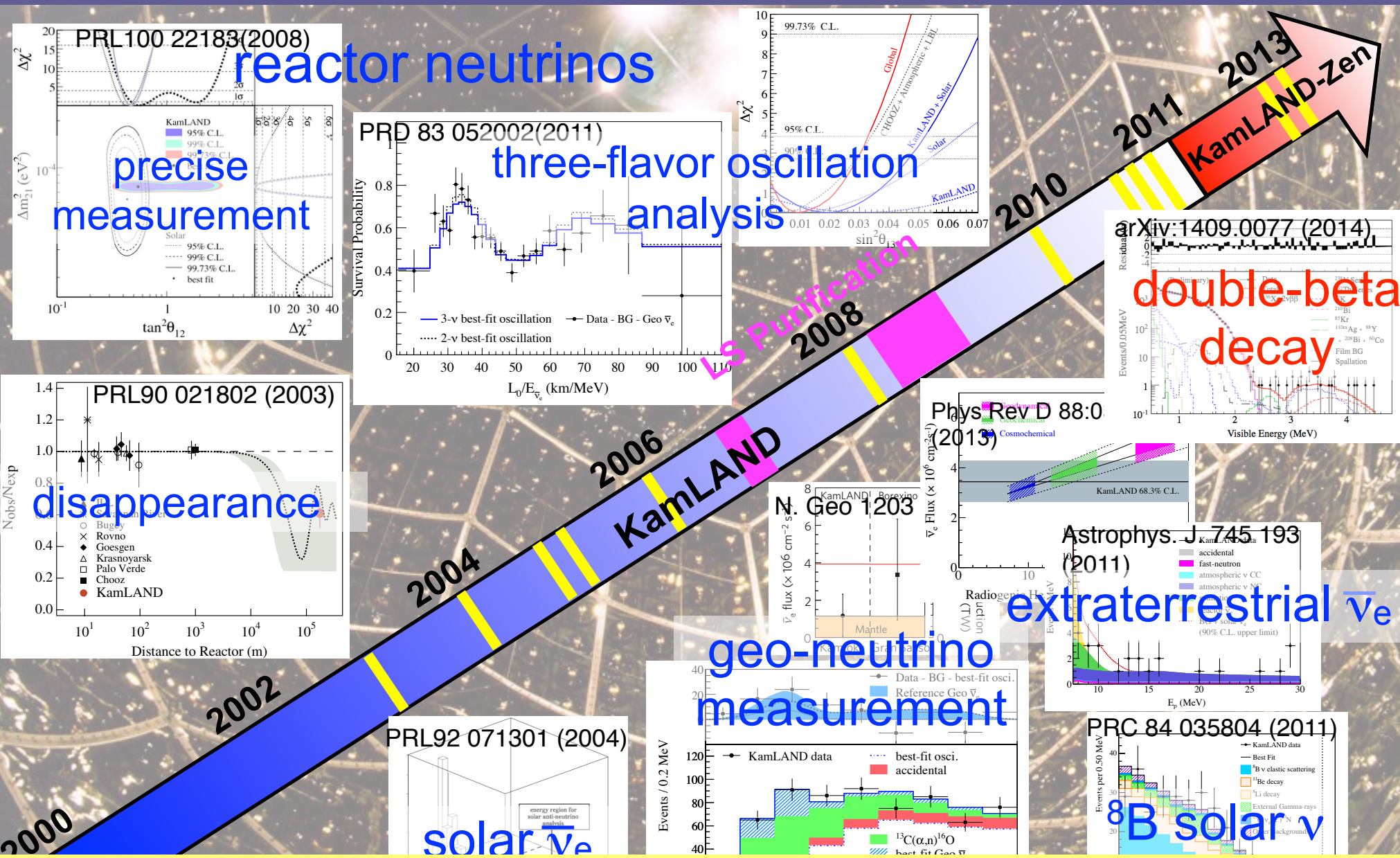
- * Dodecane (80%) Pseudocumene (20%) PPO (1.36 g/l)
- * extremely low impurity ($^{238}\text{U}:3.5 \times 10^{-18}\text{g/g}$, $^{232}\text{Th}:5.2 \times 10^{-17}\text{g/g}$)

1,325 17inch + 554 20inch PMTs

* Photo coverage 34%

Water Cherenkov Outer Detector

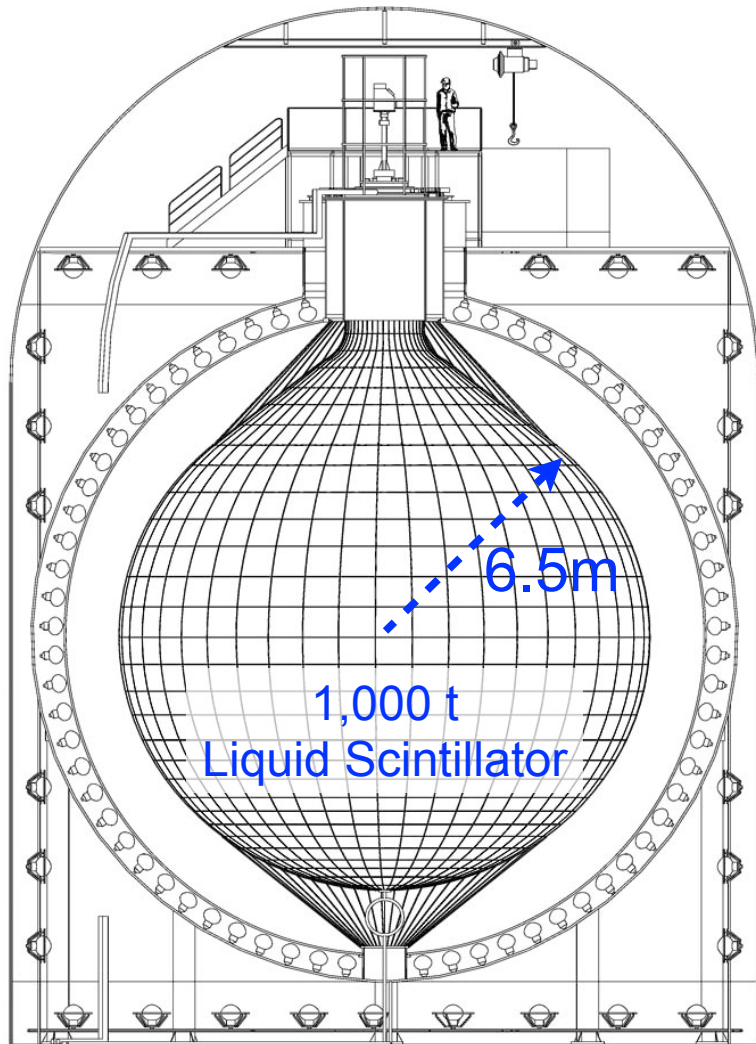
* Muon veto



We continue to study neutrino physics with KamLAND



KamLAND
2000~

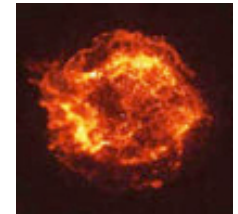
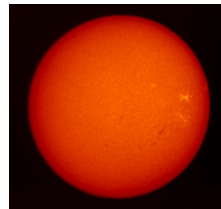
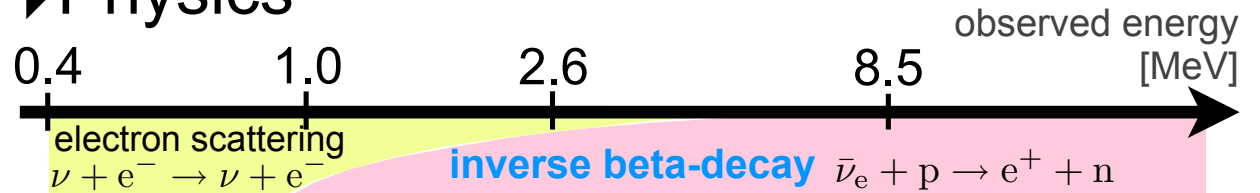


▶ Detector Features

1,000t ultra-pure liquid scintillator

^{232}U : 3.5×10^{-18} g/g, ^{238}Th : 5.2×10^{-17} g/g

▶ Physics



solar neutrinos

PRC 84, 035804 (2011)

geo neutrinos

Nature Vol. 436 (2005)
Nature Geoscience 4, 647-651 (2011)

reactor neutrinos

PRL 100, 221803 (2008)
PRD 83, 052002 (2011)

supernova neutrinos, etc.

PRL 92, 071301 (2004)
Astrophys. J. 745, 193 (2011)

Different neutrino physics in a wide energy range



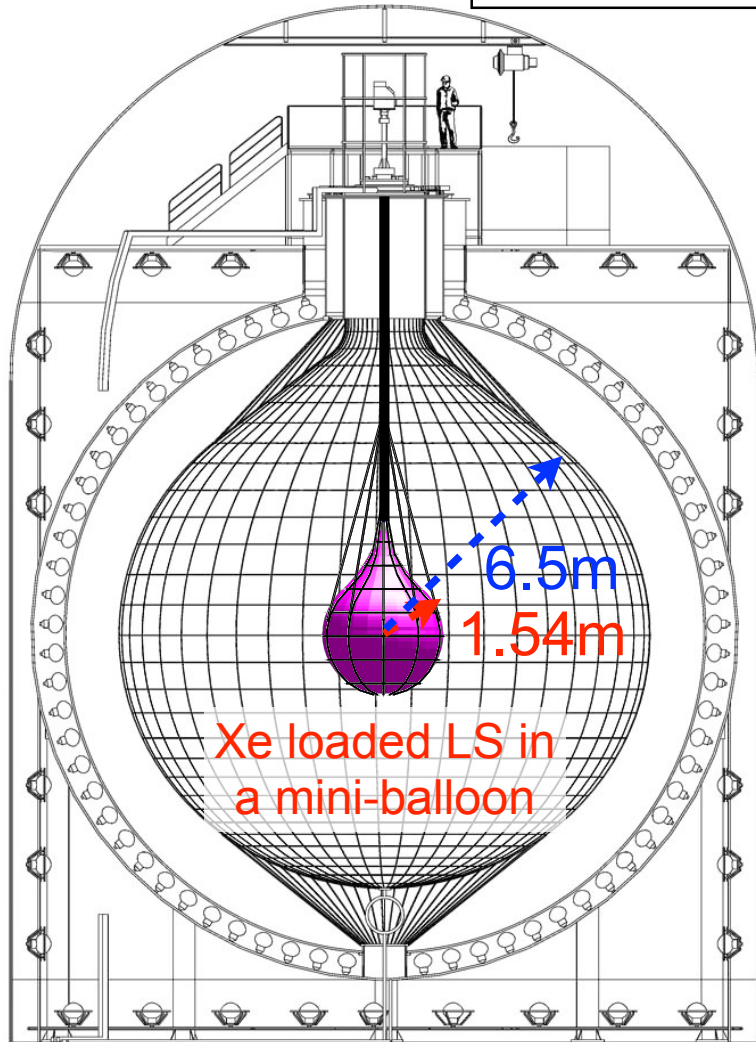
▶ Detector Features

KamLAND-Zen

2011~

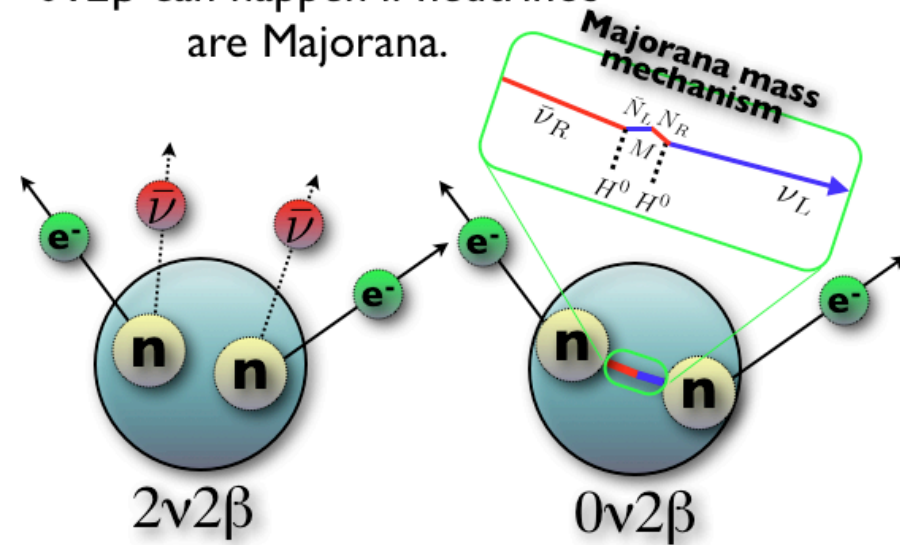
Zero Neutrino
double beta decay search

^{136}Xe loaded LS was installed in KamLAND
(383 kg of ^{136}Xe enriched Xe installed)



▶ Physics

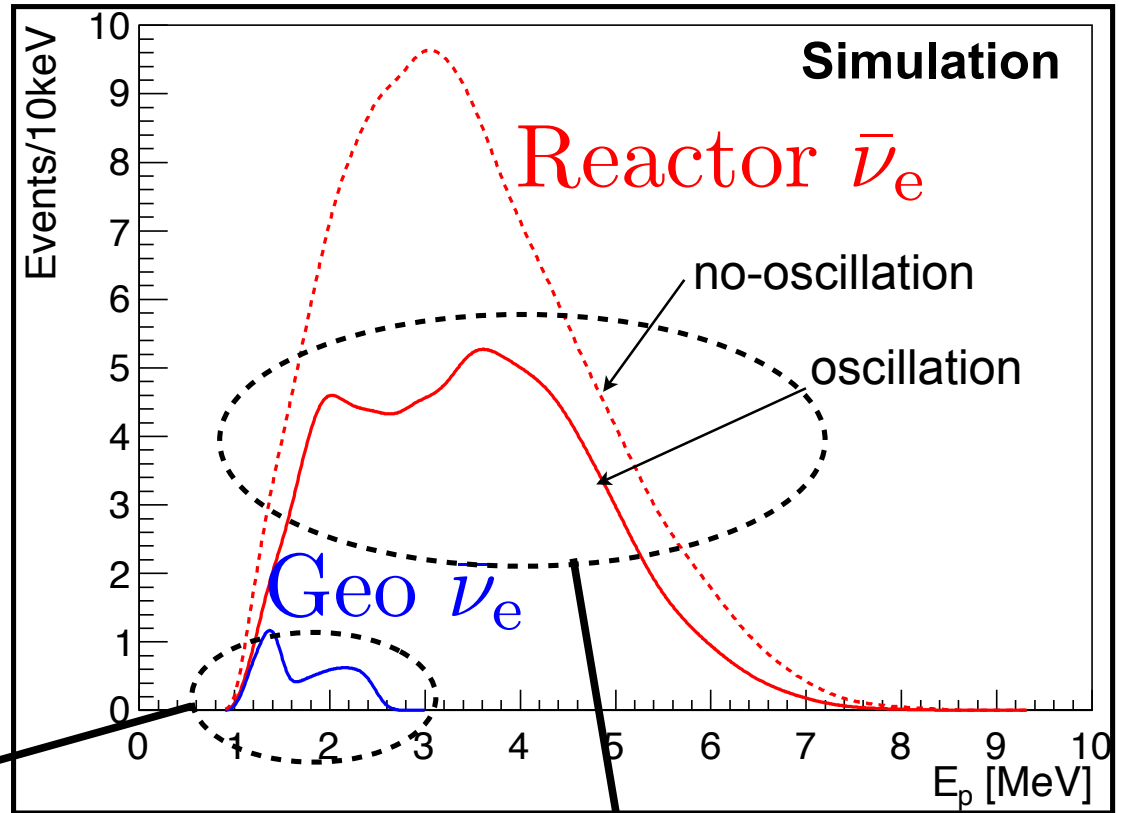
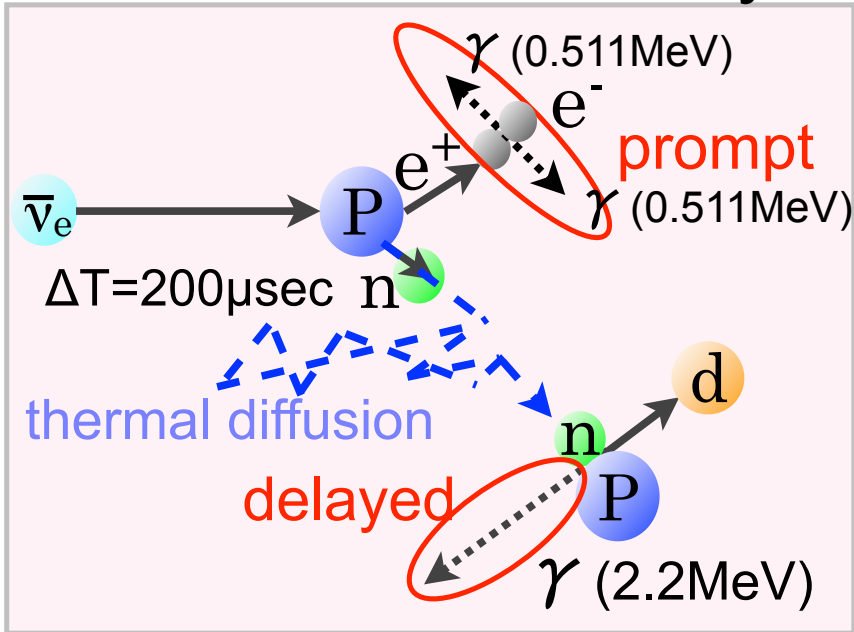
$0\nu 2\beta$ can happen if neutrinos are Majorana.



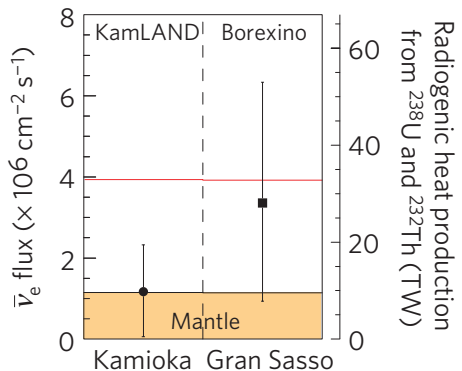
neutrino-less double beta decay

Continue to use LS volume outside of mini-balloon to measure anti-neutrino signals

inverse-beta decay

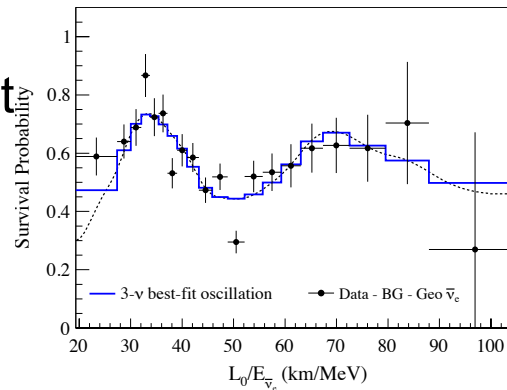


Geoneutrinos : Neutrino Application



- Direct measurement of radiogenic heat contribution

Neutrino Property Study

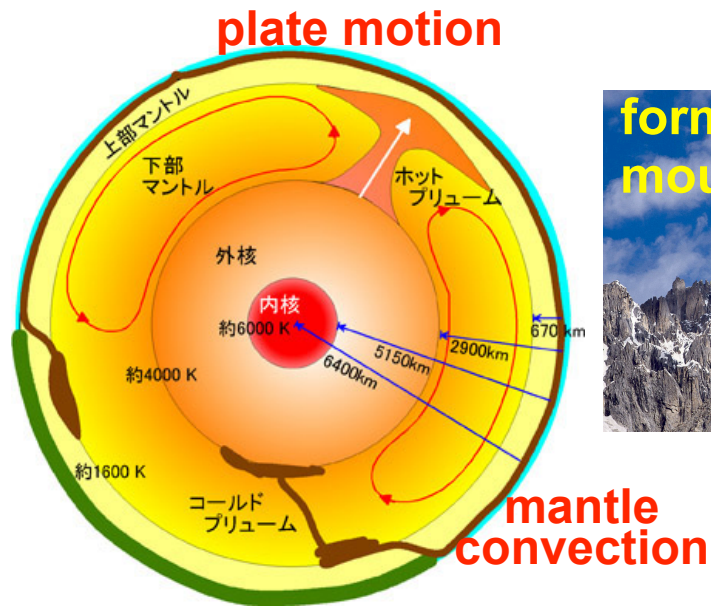


- Signature of neutrino oscillation

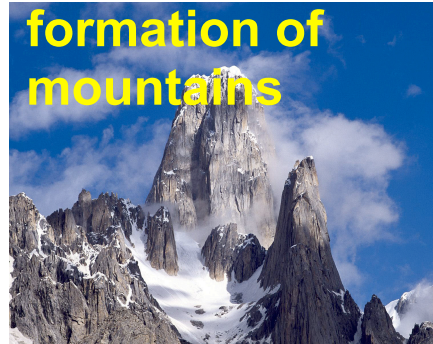
- Precise measurement of oscillation parameters

Contents

1. Introduction
- 2. Geo-neutrino Measurement Results**
3. Future Prospects
4. Summary



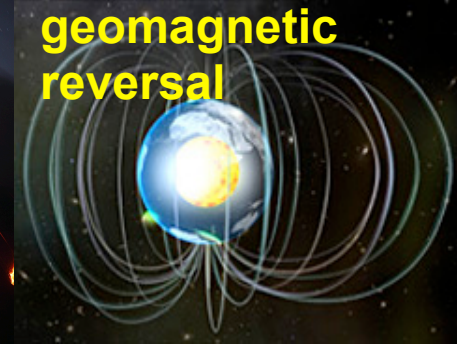
formation of mountains



earthquake
•volcano



geomagnetic reversal



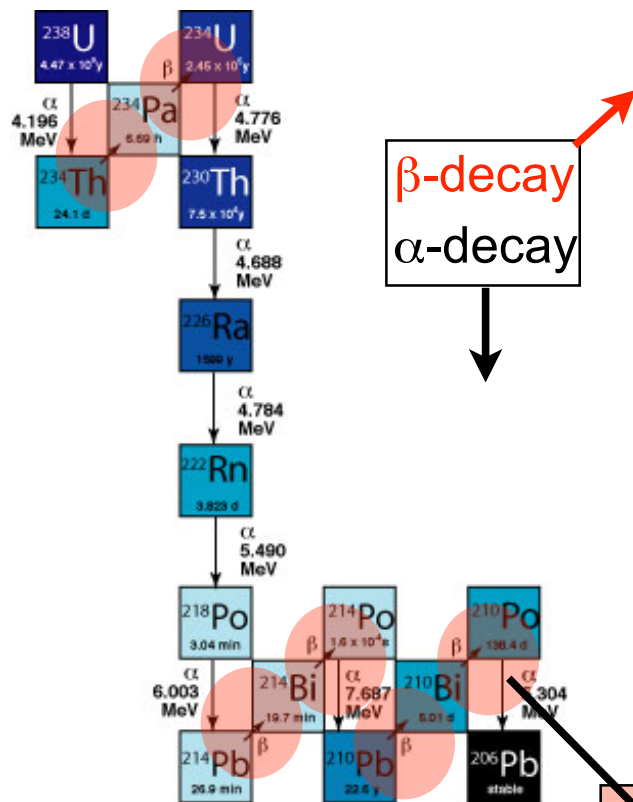
Question on geophysical activity

- What are energy sources? How much energy?
- How is the mantle convecting, single or multi-layer convection?
- Why is the frequency of geomagnetic reversals random?

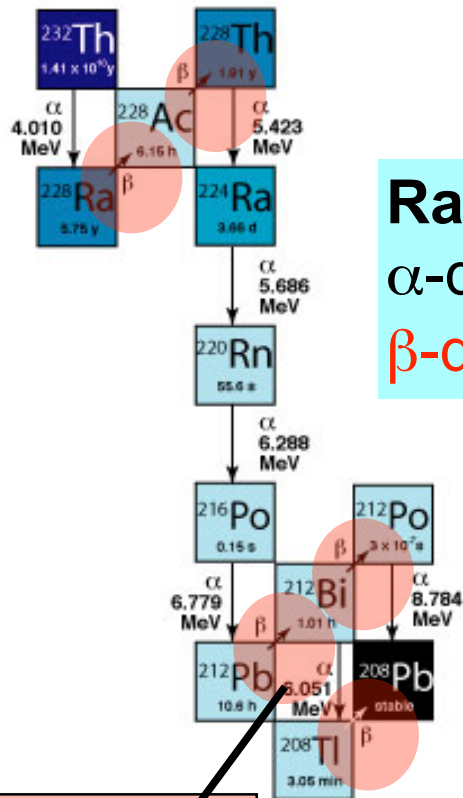
→ It is important to find out the terrestrial heat.

- (1) Radiogenic heat from **U, Th, K decay**
- (2) Release of gravitational energy through accretion or metallic core separation
- (3) Latent heat from the growth of inner core

U-series



Th-series



β -decay
 α -decay

Radiogenic heat :
 α -decay or
 β -decay emitting “anti-neutrinos”

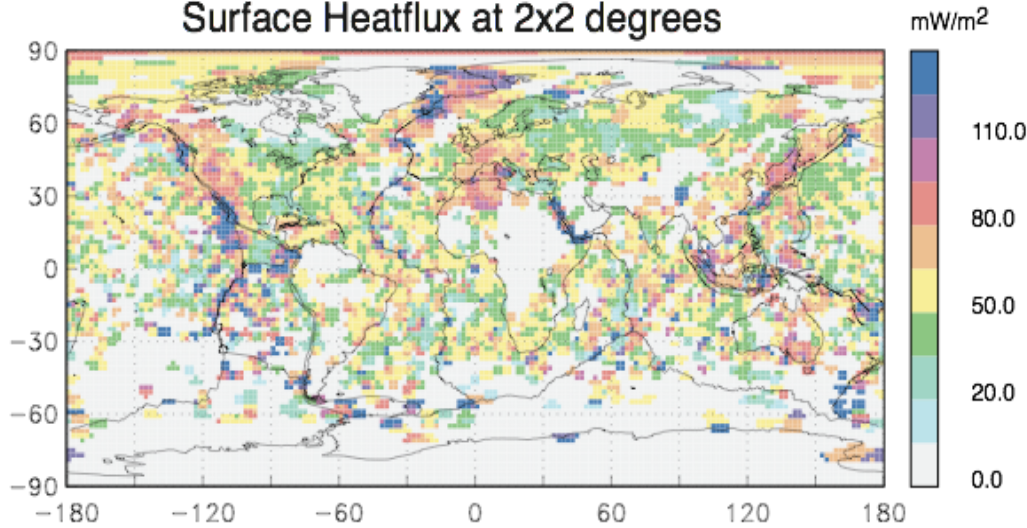
anti-neutrino
from β -decay

☑ Surface heat flow

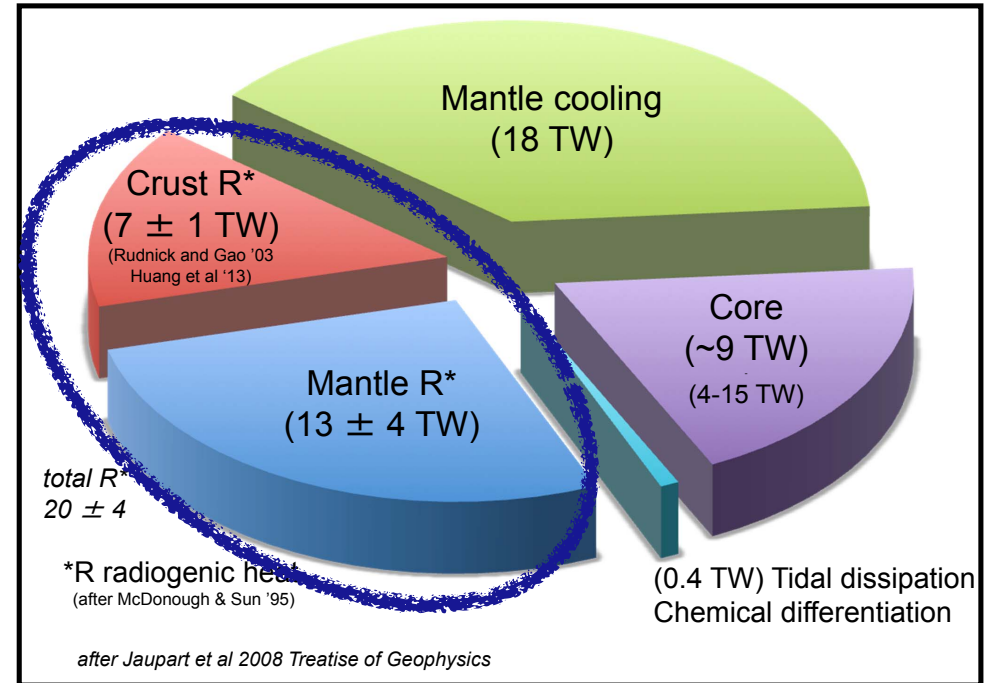
46 ± 3 TW

crust heat flux measurement & calculation

Surface Heatflux at 2x2 degrees



Pollack et al. 1933, Rev. of Geophys.



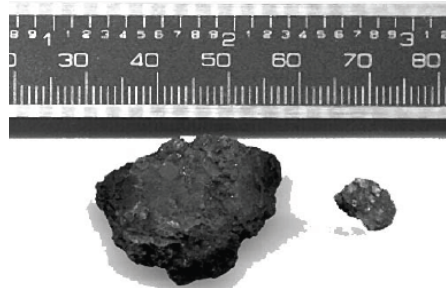
Almost half of radiogenic heat contributes to the surface heat flow.

☑ Radiogenic heat in the Earth

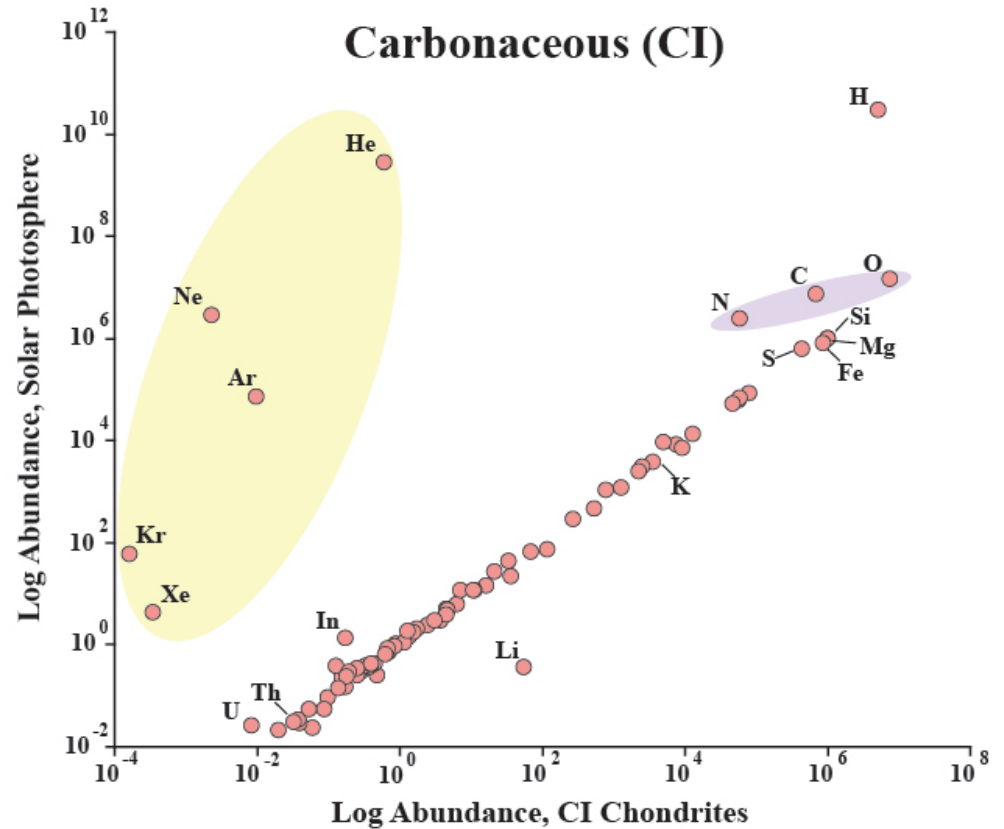
10~30 TW

Bulk Silicate Earth (BSE) model
composition of chondrite meteorite

U : 8 TW
Th : 8 TW
K : 3 TW

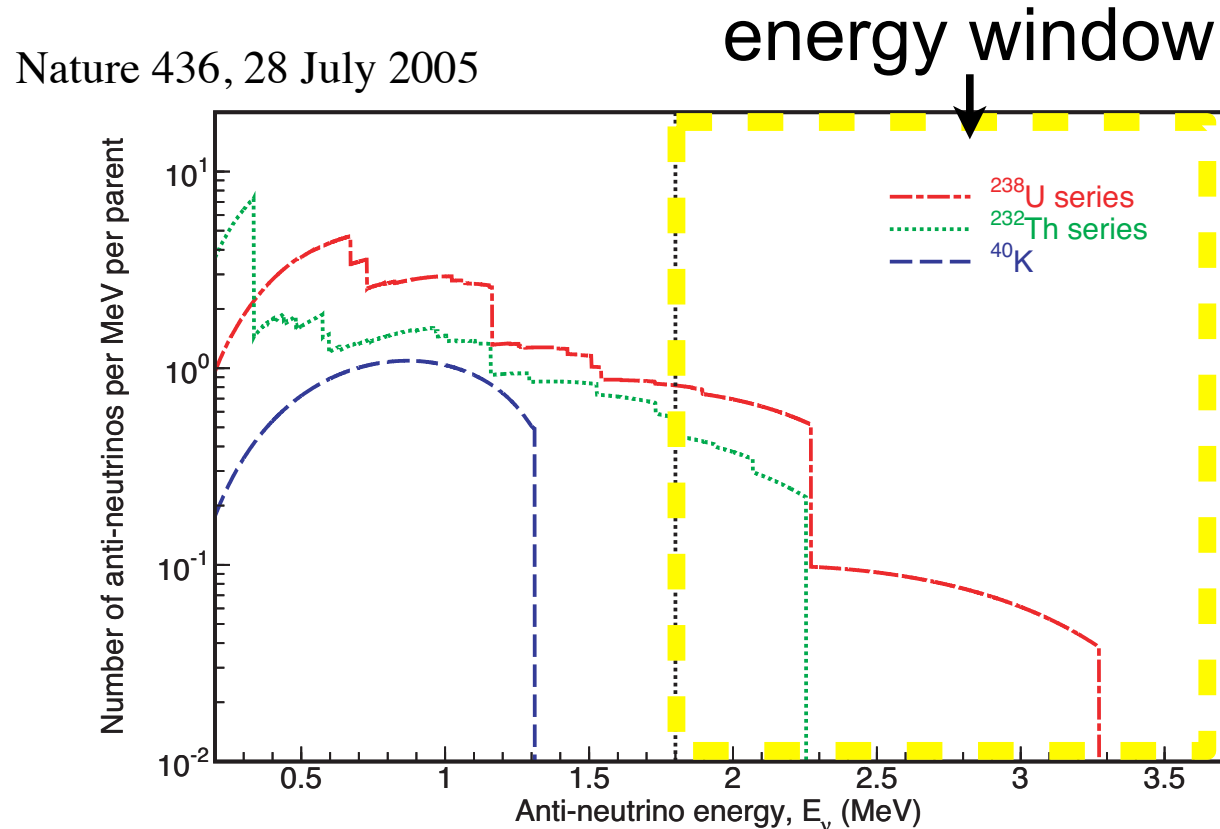


This is not “direct measurement”.

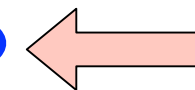
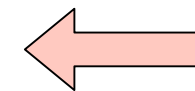
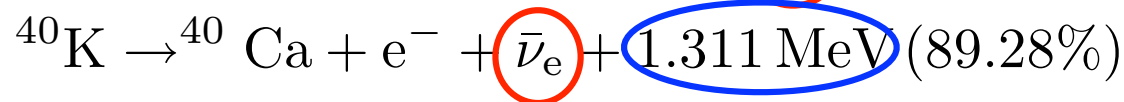
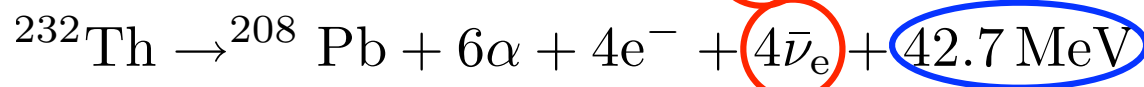
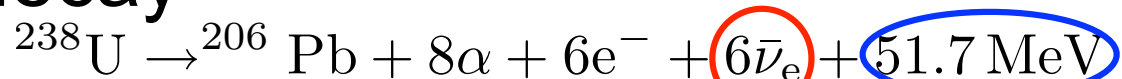


Geo-neutrino can directly test radiogenic heat production.

Geo-neutrinos are a unique, direct window into the interior of the Earth!

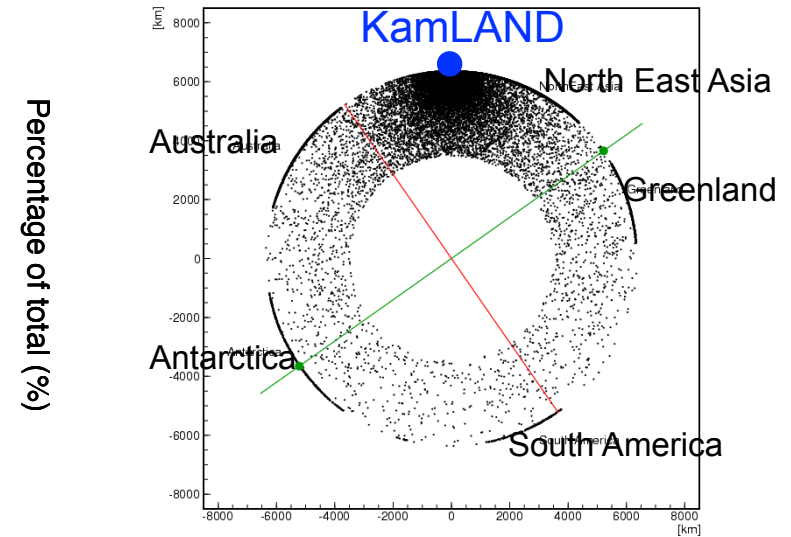
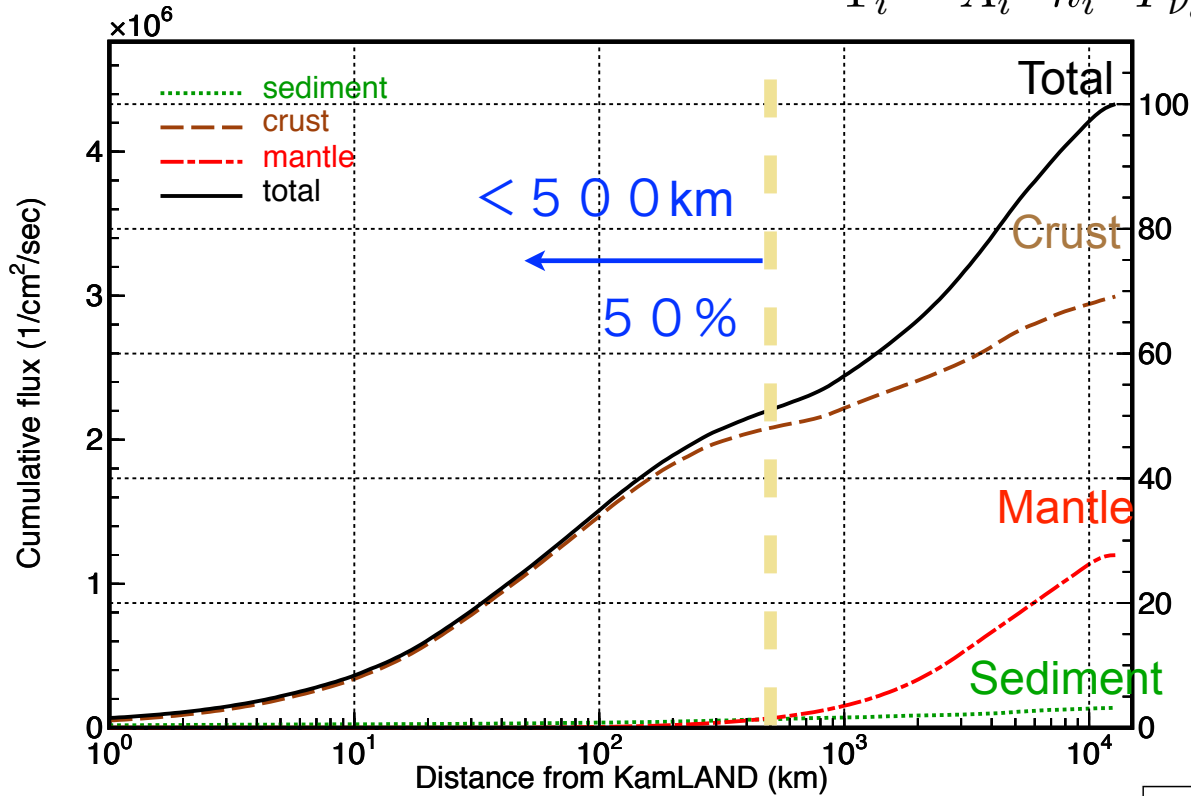


beta-decay

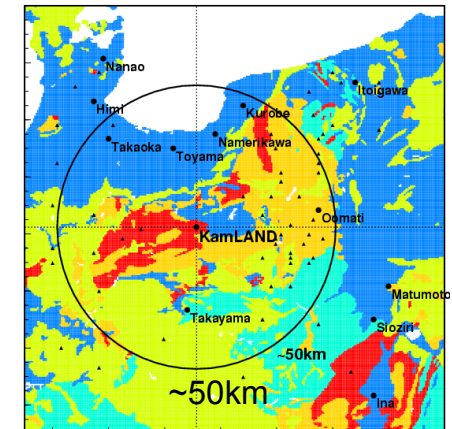
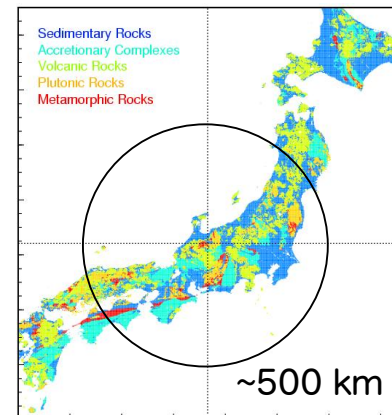


detectable

$$\Phi_i = A_i \cdot n_i \cdot P_{\bar{\nu}_e \rightarrow \bar{\nu}_e}(E_\nu, |\vec{L}|) \cdot \int_V \frac{a_i(\vec{L}) \cdot \rho_i(\vec{L})}{4\pi |\vec{L}|^2} \cdot dV$$

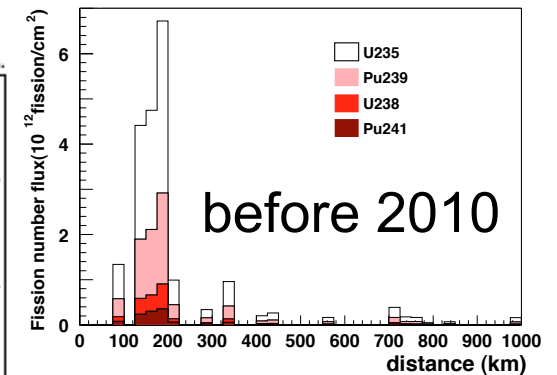
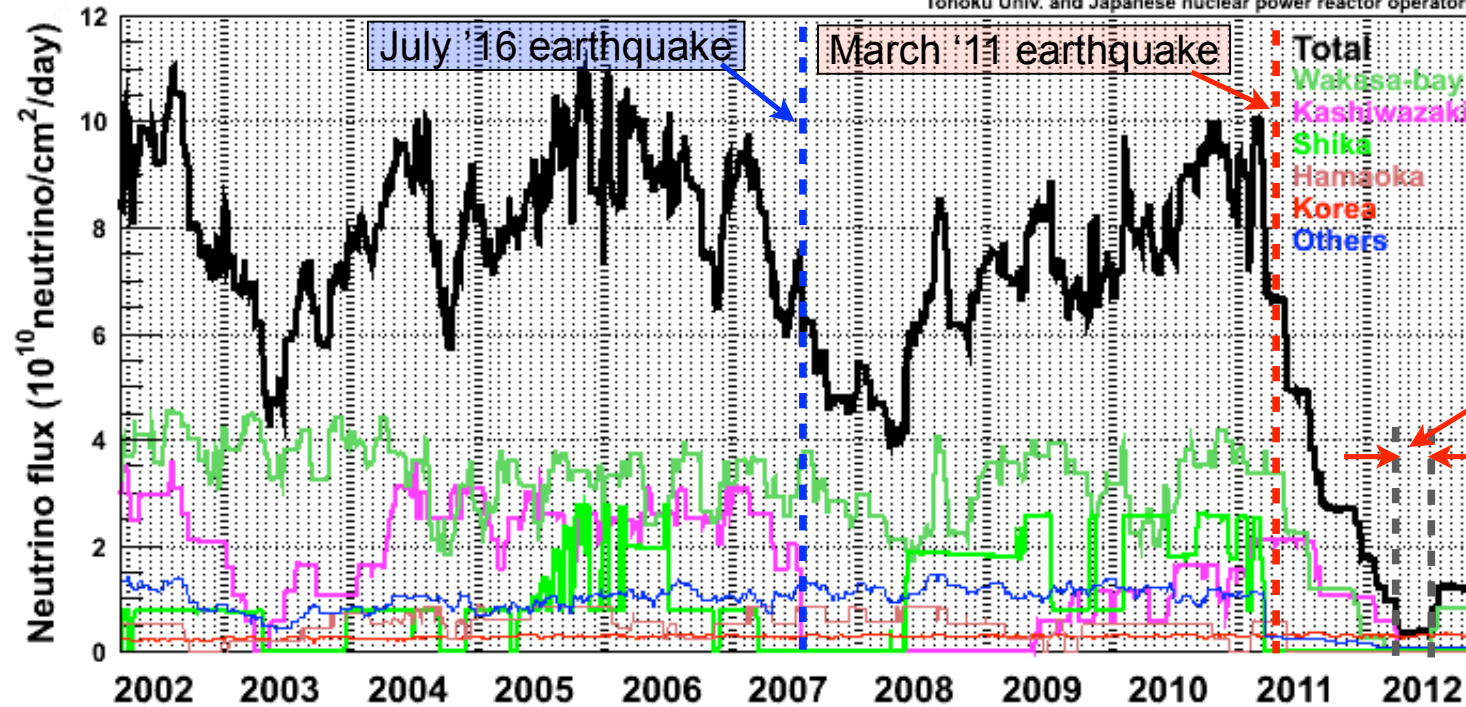


- 50%: distance < 500km
- 25%: distance < 50km
- 1~2%: from Kamioka mine

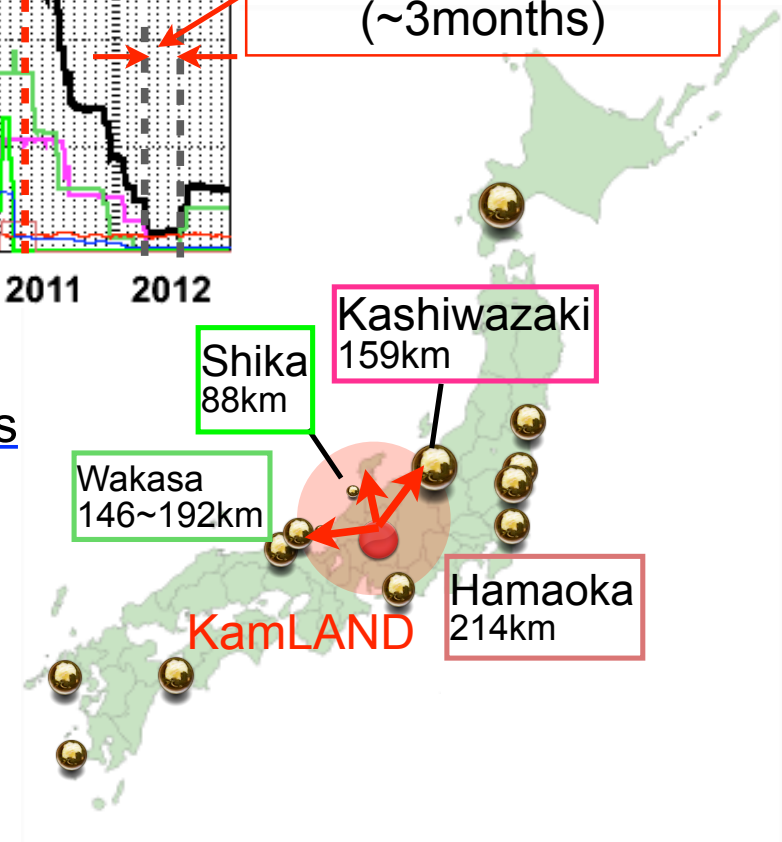


time variation of neutrino flux

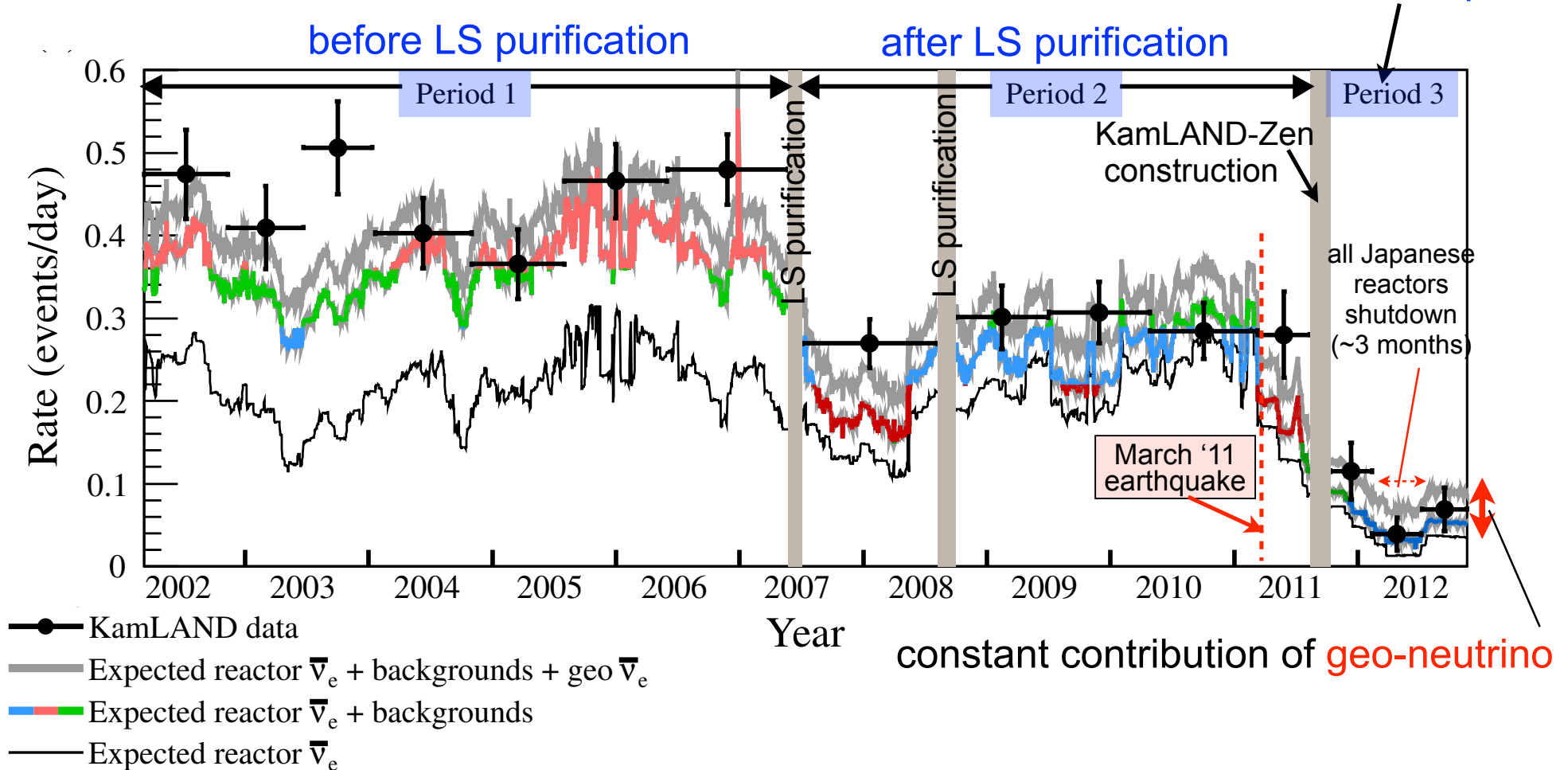
Data provided according to the special agreements between Tohoku Univ. and Japanese nuclear power reactor operators.



- Following the Fukushima nuclear accident in March 2011, the entire Japanese nuclear reactor industry has been subjected to protected shutdown.
- Reactor neutrino flux, which is outside the control of the experiment, was significantly reduced.
- This situation allows for a “reactor on-off” study of **backgrounds for KamLAND neutrino oscillation and geoneutrino analysis.**



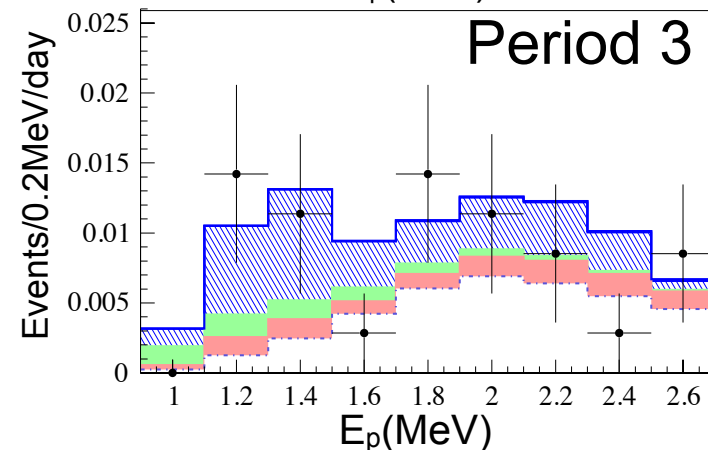
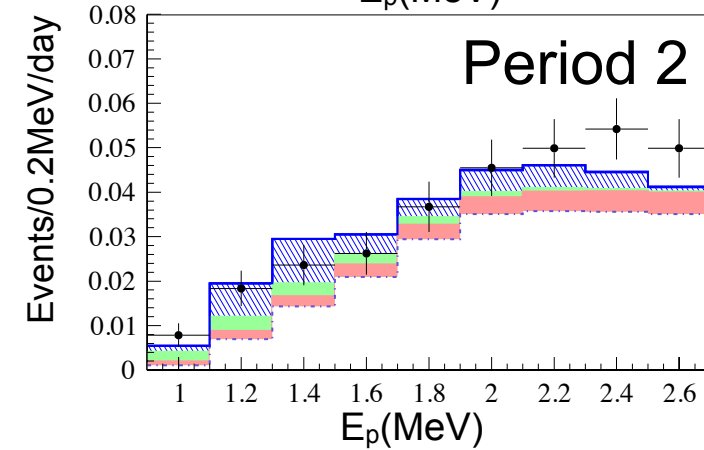
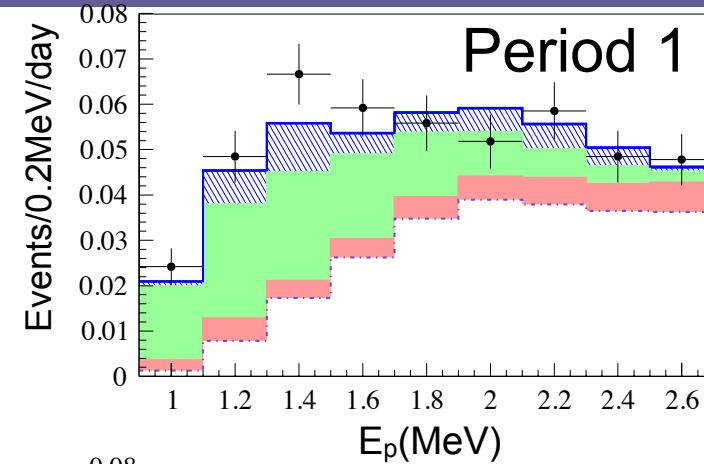
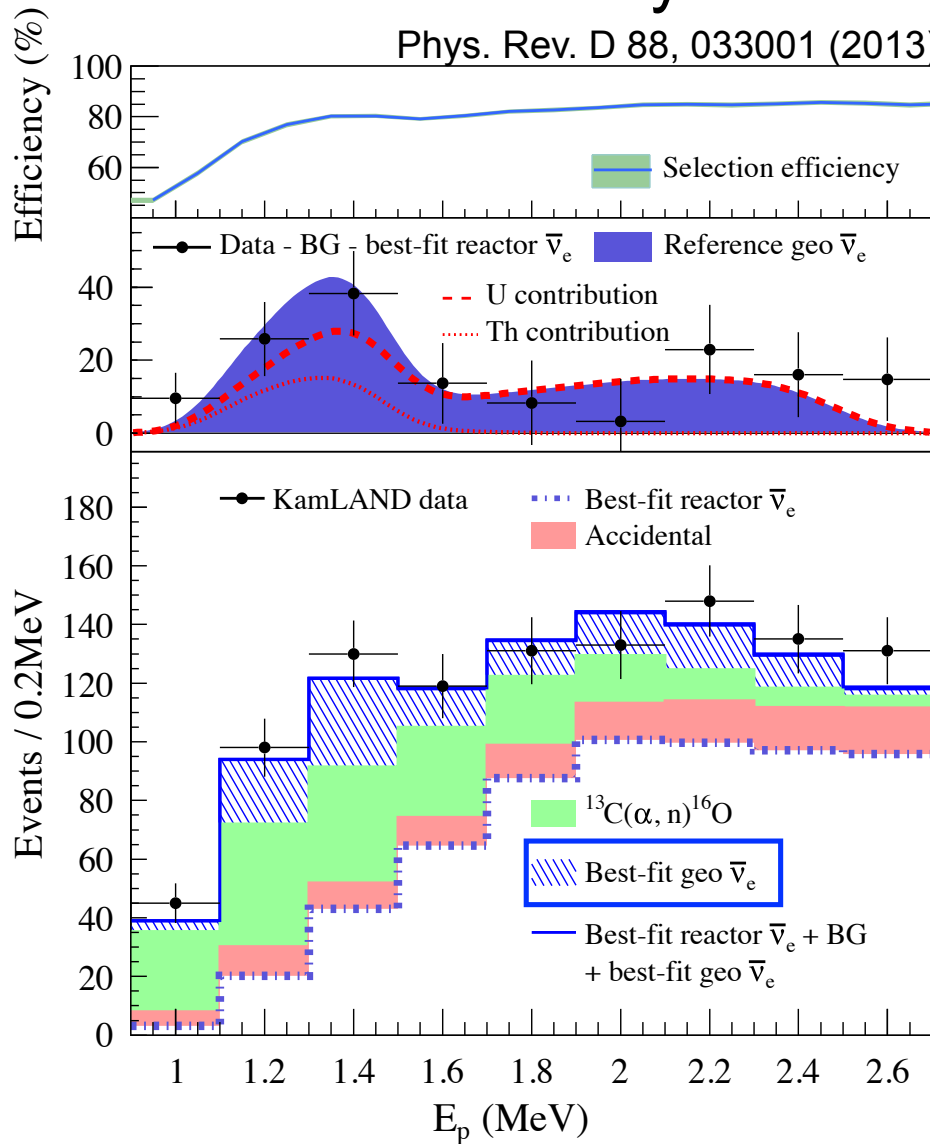
after KamLAND-Zen start, low reactor phase



- Backgrounds :
 - LS purification → non-neutrino backgrounds reduction
 - Earthquake → reactor neutrino reduction
- Constant contribution of geo-neutrino
 - Time information is useful to extract the geo-neutrino signal

data set : 2991.1 days

Phys. Rev. D 88, 033001 (2013)

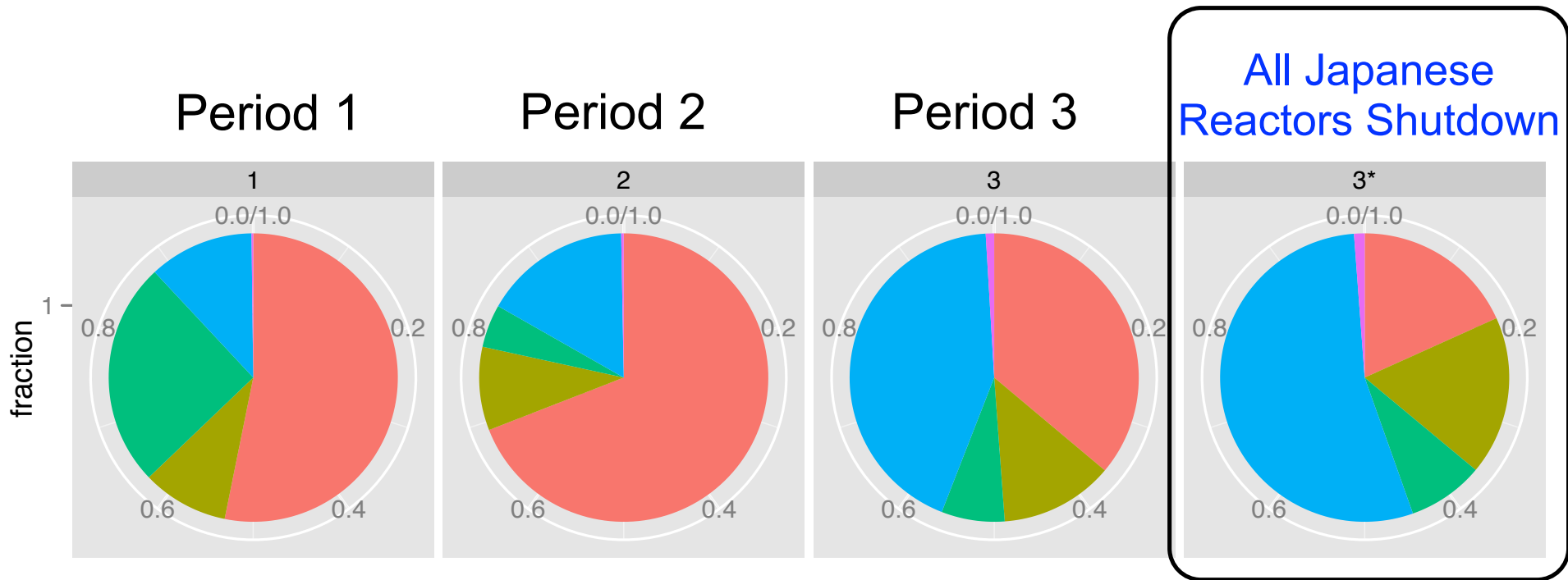


$^{13}\text{C}(\alpha, n)^{16}\text{O}$

↓ decreased

Reactor- $\bar{\nu}$

↓ decreased



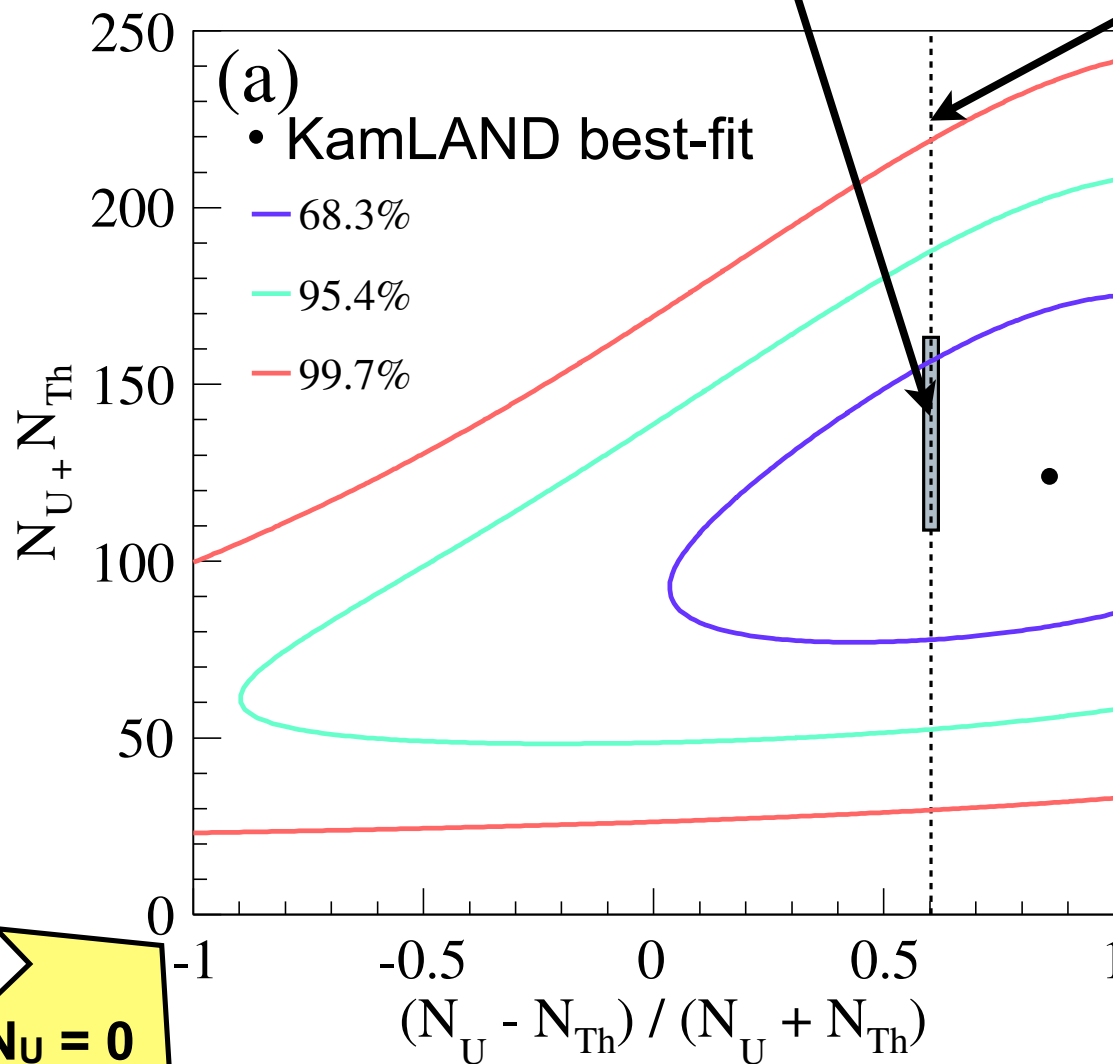
- reactor neutrino
- accidental background
- (α,n) background
- geo-neutrino
- other events

ratio between geo-neutrino signals and backgrounds is higher than 1

y-axis : number of U and Th geo-neutrino

Earth model prediction
EPSL 258, 147 (2007)

Th/U = 3.9 fixed



◆ limits on Th/U ratio

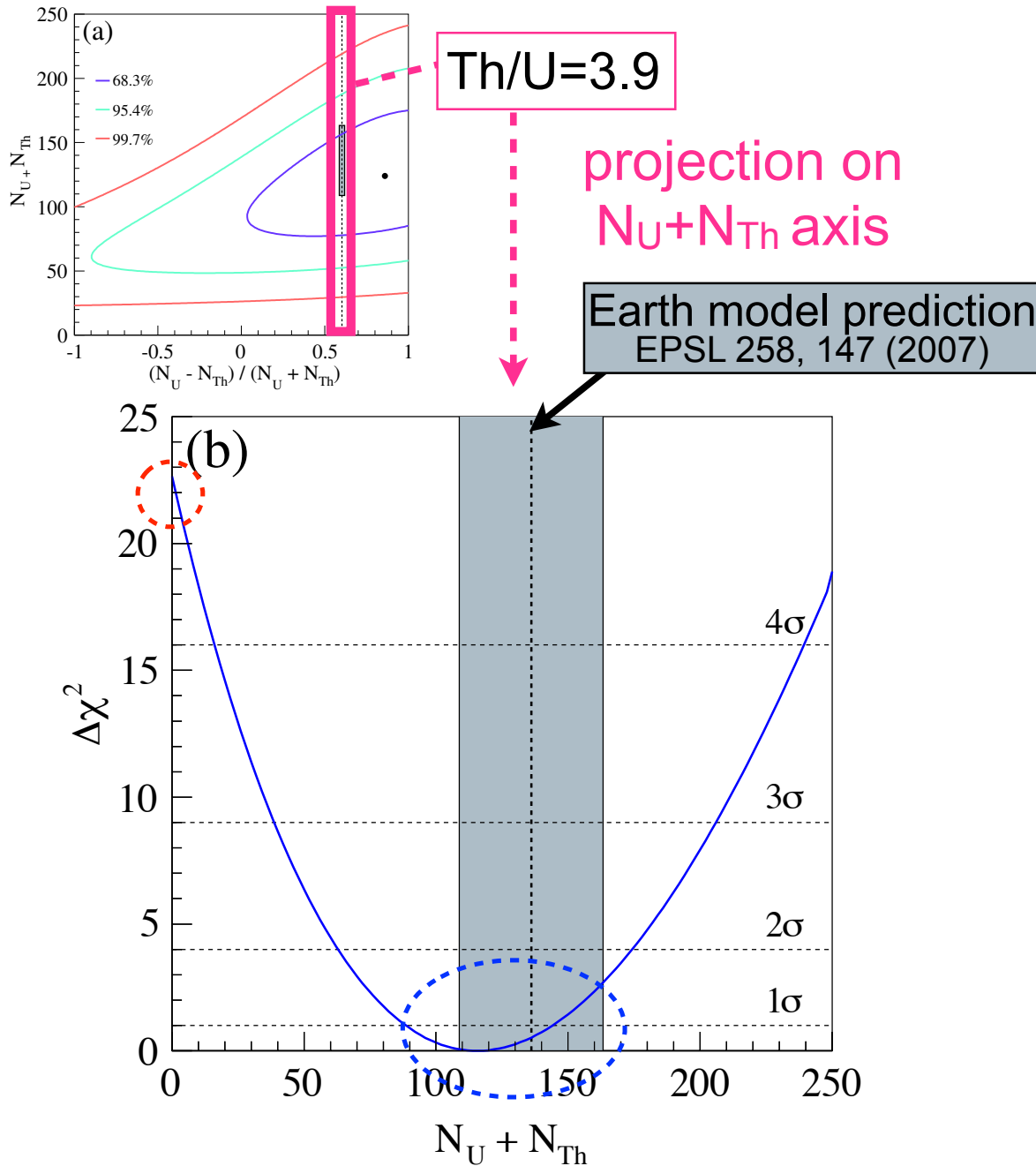
Th/U < 19 (90% C.L.)

◆ Th/U mass ratio
(Th/U = 3.9)

Number of geo-neutrino

$N_{geo} = 116^{+28}_{-27}$ events

$F_{geo} = 3.4^{+0.8}_{-0.8} \times 10^6 / \text{cm}^2 / \text{sec}$



◆ limits on Th/U ratio

Th/U < 19 (90% C.L.)

◆ Th/U mass ratio
(Th/U = 3.9)

Number of geo-neutrino

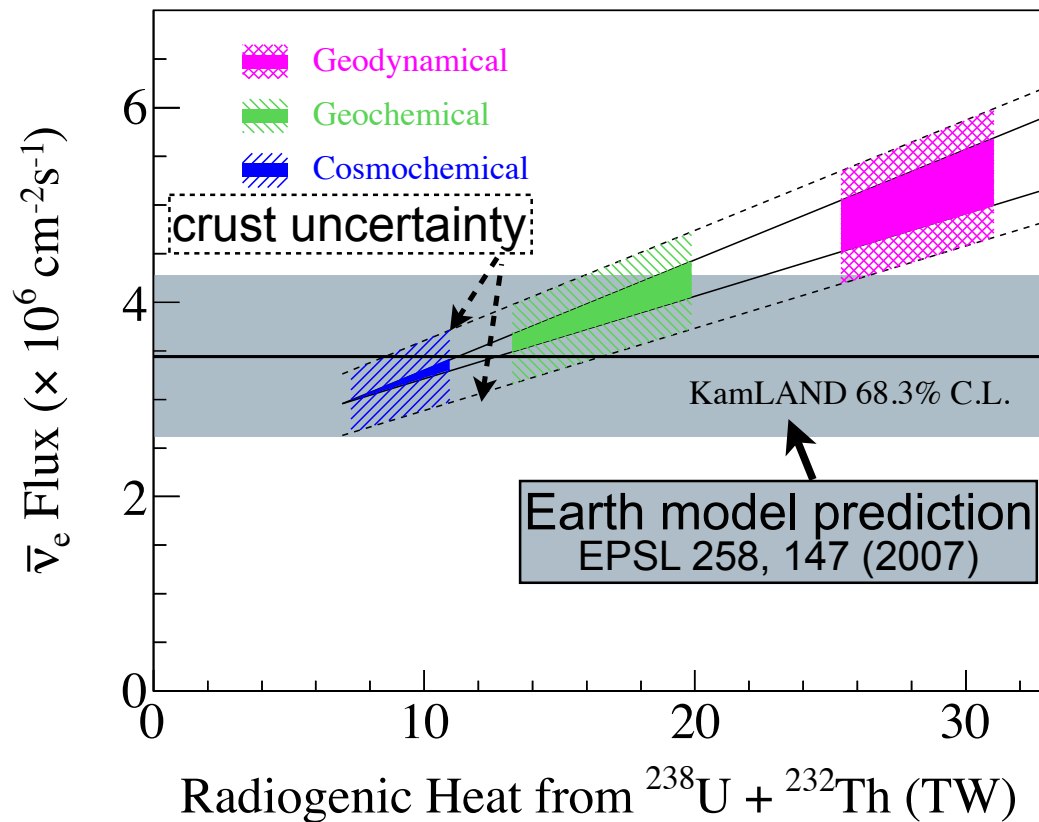
$N_{geo} = 116^{+28}_{-27}$ events

$F_{geo} = 3.4^{+0.8}_{-0.8} \times 10^6 / \text{cm}^2 / \text{sec}$

almost same as model prediction

0 signal is rejected at

99.9998% C.L.



[BSE composition models]

Geodynamical 30TW

based on balancing mantle viscosity and heat dissipation

Geochemical 20TW

based on mantle samples compared with chondrites

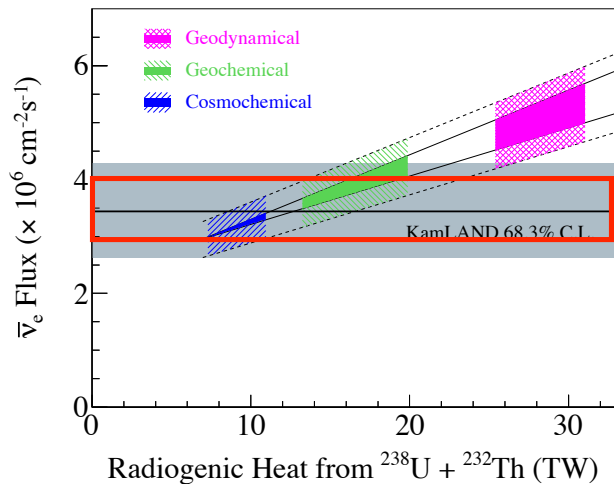
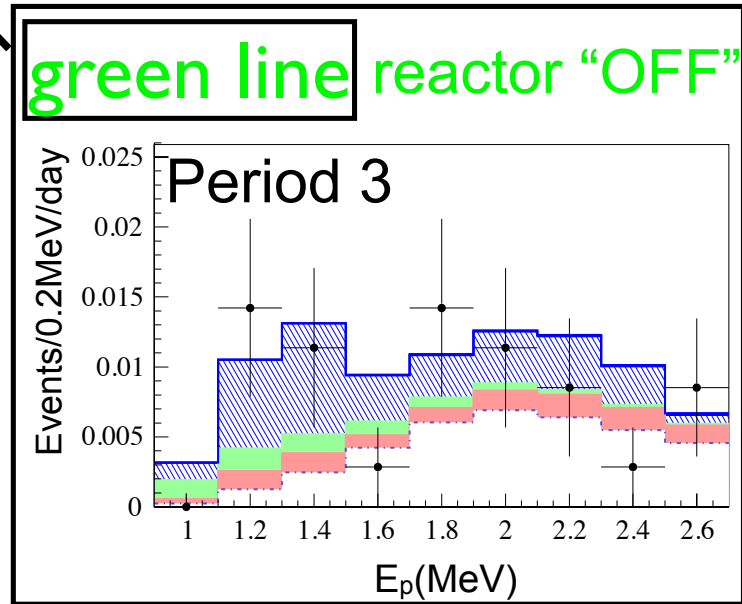
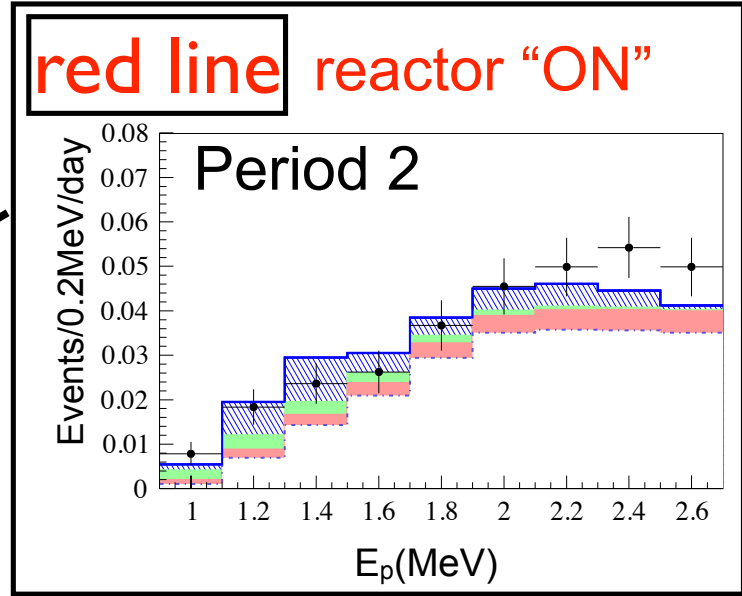
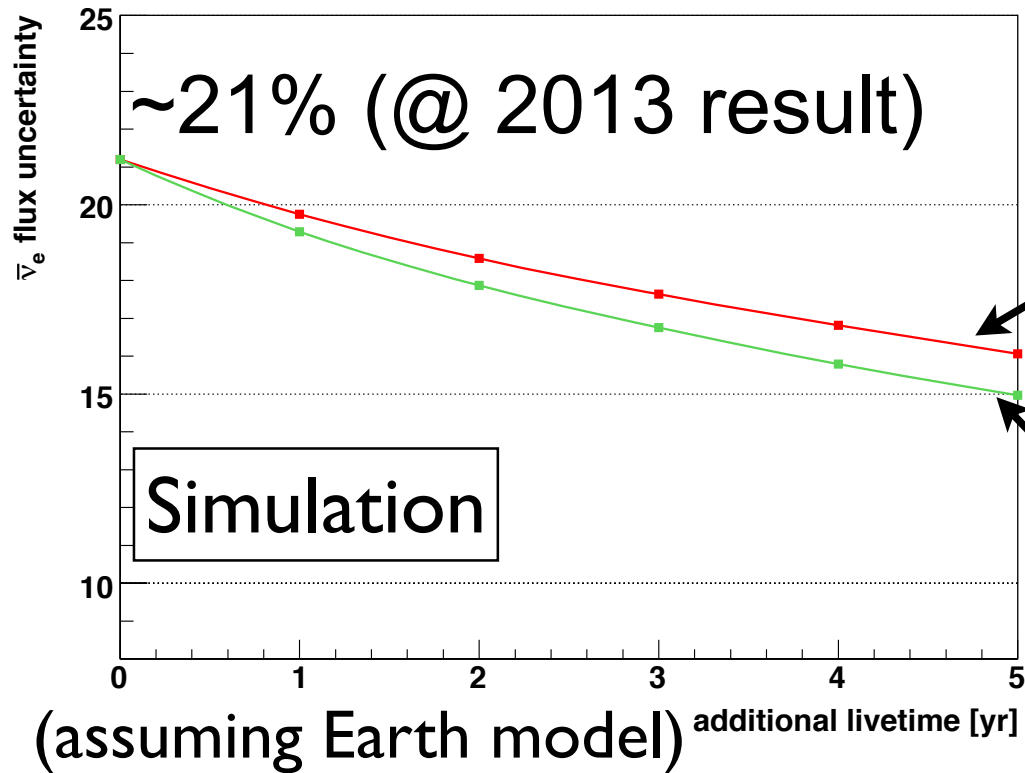
Cosmochemical 10TW

based on isotope constraints and chondritic models

- KamLAND geo-neutrino flux translates to a total radiogenic heat production : **11.2^{+7.9}_{-5.1} TW**
- The geodynamical prediction with the homogeneous hypothesis is disfavored at **89% C.L.**
- All BSE compositional models are still consistent within **~2 σ**.

Contents

1. Introduction
2. Geo-neutrino Measurement Results
- 3. Future Prospects**
4. Summary



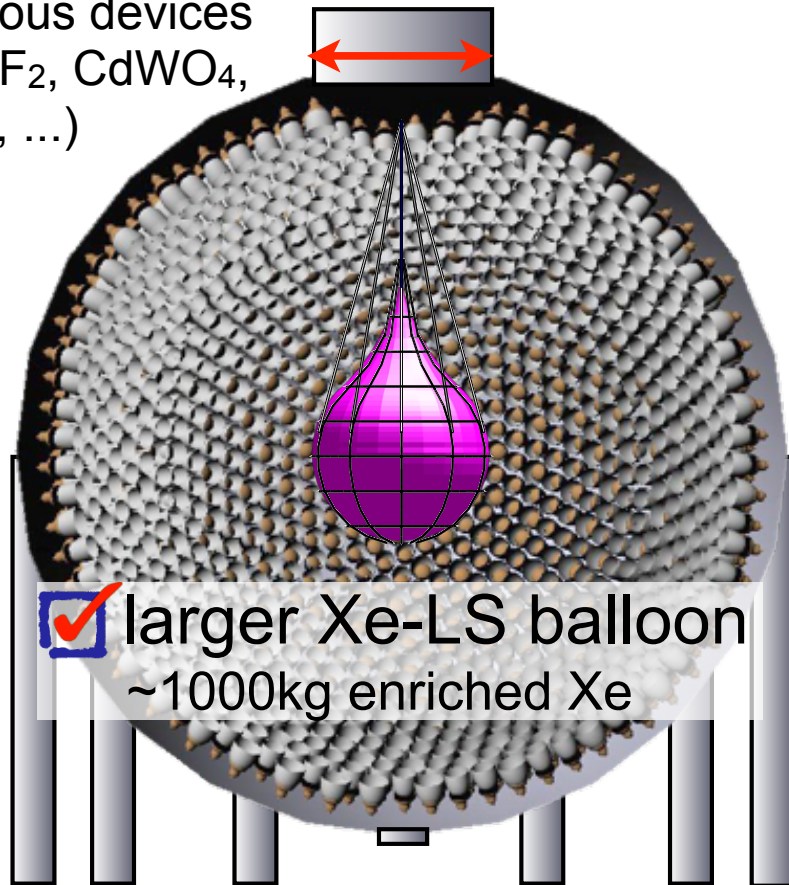
We will achieve
 15~16% uncertainty
 with additional 5 year
 measurement.
 (We already have
 another 2-year data.)

▶ KamLAND2(-Zen) : better energy resolution 21/23

upgrade to KamLAND

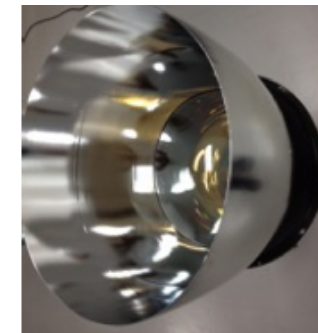
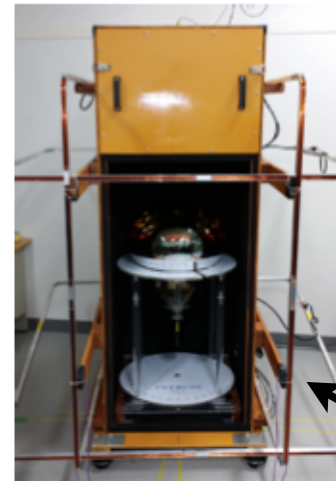
☑ enlarge opening

accommodate
various devices
(CaF₂, CdWO₄,
NaI, ...)



☑ High performance

* High Q.E. PMT * Winstone Cone



← 17"Φ → 20"Φ,
ε = 22% → 30%

Photo-coverage > ×2

Light Collecting Eff. > ×1.8

* New Liquid Scintillator

LAB based LS (8,000 → 12,000 photon/MeV)

energy resolution improved

: 6.4%/√E[MeV] → 4.0%/√E[MeV]

photon
yield

×1.9

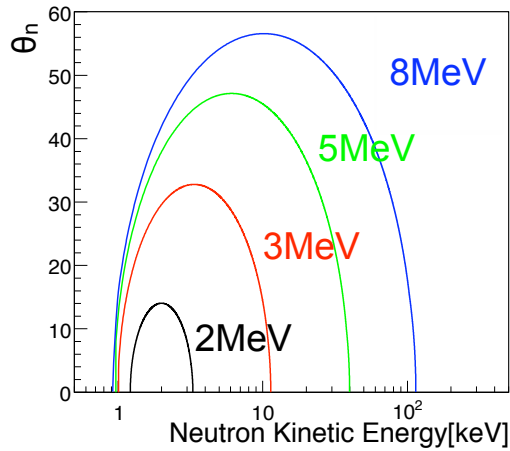
×1.4

geo-neutrino measurement

* improvement of U/Th ratio

* fiducial volume enlargement

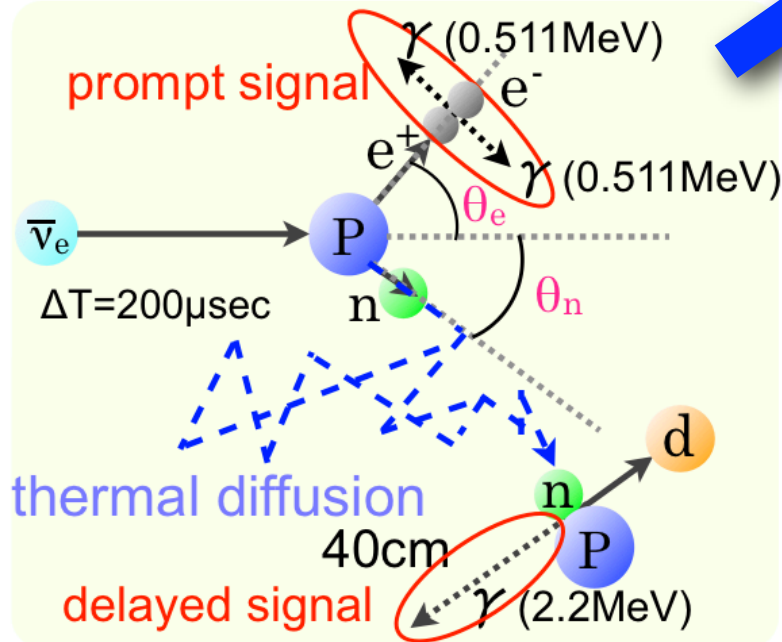
► Directional measurement see also "Future Projects" session 22/23



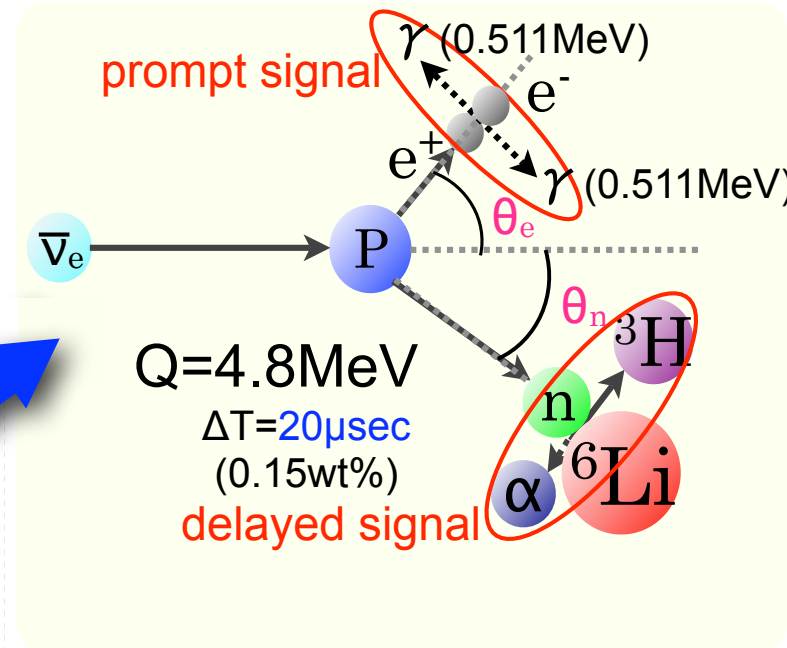
neutron has directional information of anti-neutrino

$$E_{\bar{\nu}_e} < 3\text{MeV} \rightarrow \theta_n < 35^\circ$$

[current liquid scintillator]



[Li loaded liquid scintillator]



- large neutron capture cross section (${}^6\text{Li}$ 940 barns vs ${}^1\text{H}$ 0.3 barns)
- α doesn't travel far

+

high vertex resolution imaging detector

- higher than 2 cm resolution (PMT $\sim 10\text{cm}$)

Contents

1. Introduction
2. Geo-neutrino Measurement Results
3. Future Prospects
- 4. Summary**

▶ The KamLAND experiment measures anti-neutrino from various sources over a wide energy range

▶ **Geo-neutrino**

- Observed flux is fully consistent with Earth models

- Results for low reactor background:

Geo-neutrino observation is very sensitive

- Now we enter the era of conducting critical tests of Earth models

▶ It is important for “Neutrino Geoscience” to further connections between geoscience and neutrino physics.