Bubble Chambers for Dark Matter Detection:

COUPP (Chicagoland Observatory for Underground Particle Physics)









The COUPP Team



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Wish List for a Tough Problem



Dream Detector:

- Scalable
- Cheap
- Radiopure
- Good backgroundrejection
- Multiple active nuclei
- Easy to swap targets

A bubble chamber (bulk superheated liquid detector) potentially fulfills all those wishes...if it actually works

Bubble Chambers for Rare Event Searches?

- In recent HEP, bubble chambers with metallic parts were stable for ~ 10 milliseconds...killer dead time
- But the many of the original chambers were glass
- Condensed matter research strongly suggests that both bulk and surface spontaneous nucleation rates can be made low enough





Fig. 2. Assembled bubble chamber and jacke

Roger Hildebrand at UofC holding the first Hydrogen bubble chamber

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Try This At Home?



A single microscopic interaction (nuclear recoil) gives a macroscopic signal (boiling)

→ The WIMP signal is just as large as the fork signal!

β and γ Insensitivity

- Most experiments, after heroic cleaning and shielding, still must cut many β , γ events
- Bubble chambers can be operated essentially MIP-blind



Freedom to Choose Target

In general SUSY, channel of first detection is unknown:

- Spin-dependent: use fluorine
- Spin-independent: use something heavy

We use $CF_{3}I$, for excellent sensitivity in both channels.

Other advantages of CF₃I :

- Inexpensive (40 USD / kg)
- Near room temperature operation (~ 40 C)
- Safe (non-toxic, fire extinguisher)



Recently expanded by Baltz (WMAP-II range 0.103 $\Omega h^2 < 0.112$ lambda.gsfc.nasa.gov/product/map/current/parameters.cfm.) Courtesy of P. Gondolo

Freedom to Swap Target



Dark blue: LSP (SUGRA) Light blue: LKP

(Bertone, Cerdeno, Collar, and Odom, to appear in PRL, arXiv:0705.2502)

Swapping Target Liquids:

• Upon discovery, makes ~ orthogonal cuts in parameter space

• Alternatively, use different targets for neutron rejection. (Response of CF3I and C3F8 to neutrons is very similar, but, for example, SI WIMP response is very different)

2 kg Prototype Chamber



Bulk CF3I touches only the quartz



Movie posted at http://kicp.uchicago.edu/~odom/presentations/bubble3.avi



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Leaving "Normal" BC Mode Far Behind



Decreasing superheat

- Muon tracks are visible only at high superheats
- We operate at low superheat, MIP blind
- 1 nuclear recoil = 1 bubble

Stereo View for Bubble Position



Position reconstruction is important for background rejection, of both wall events and neutron events (multiple scattering)

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Operation at 300 m.w.e., Fermilab







Demonstration of MIP-Blindness



 Intrinsic γ rejection
 better than 10⁻¹⁰ at 10
 keV threshold (best measured MIP-rejection)

- No cuts required!
- No need for β , γ shielding or attention to β , γ radiopurity. E.g. ¹⁴C are *O* (100 / kg-day)

Neutron calibration, in situ



- Switchable Am/Be source yields
 5 +/- 0.5 n /s when on and O (0.2 n/day) when off
- Left: blind absolute comparison, consistent with 100% efficiency
- Moderated spectrum produces recoils approximating WIMP spectrum

Current Wall Background: *α*-Emitters



- Wall rate (0.9 /cm²/day) OK for small vessels (200 /day for 2 kg chamber)
- γ-counting of our Ge-214 natural fused quartz found U (Th) at typical level of 50 (30) ppb consistent with 100% of wall-event rate
- Commercial α -counting gives consistent rate
- Next-generation vessels made from synthetic fused silica, with expected U, Th level < 10 ppt
- New vessels also undergo extra quartz etch before final annealing and are then sealed, to guard against embedded Rn progeny

Current Bulk Background: Radon & Progeny

Bulk background is consistent with 100% Rn and progeny α 's

- 1. Flat rate with sharp onset at predicted value
- 2. Analysis of time correlation between events supports 100% Rn & progeny, 100% efficiency



Any WIMPs in the Data?

Q: How much WIMP signal could be hiding beneath the radon & progeny?



Note: these are integrated energy spectra

• Total exposure of 250 kg-days (Dec '05 – Nov '06), with live-time fraction reaching 80%. 51 kg-livedays used for WIMP analysis.

- Dashed curves: null hypothesis function (NHF)--only Rn & progeny
- Solid lines: expected WIMP response function (WRF) for $\sigma^{SD(p)}=3$ pb
- Perform 2-parameter fit a*NHF + b*WRF (constrain b≥0). b=0 is generally favored. (Error matrix analysis, checked against MINUIT, is used to obtain 90% C.L.)
- Cross-section limits obtained from weighted mean of WRF parameter "b", from 3 data sets

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COUPP (2007) Exclusion Limits



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Under construction: 20 kg Windowless Chamber

Pressure vessel is off-the-shelf water pipe. Readout is from internal cameras and encapsulated piezos.





20 kg synthetic fused silica vessel, sealed against exposure to Rn

Under construction: 20 kg Windowless Chamber



Shielding installation at TARP tunnel (300 mwe) near UofC



Under construction: 60 kg Chamber

For installation in NUMI tunnel, Fermilab.



30 cm x 1 meter inner vessel







Conclusions

- Scalable: liquid target with simple instrumentation
- Low cost (<1000 USD/kg)
- Outstanding γ-rejection
- Single concentration: reducing α emitters to ~10⁻¹⁷ (achieved elsewhere) probes most SUSY models
- 2 kg prototype already contributes to SD(p) limits, with lower background run underway
- Larger masses coming (20 and 60 kg modules under construction)

