



# Latest Results from The MINOS Experiment

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- Precision studies of  $v_{\mu}$  disappearance.
  - Measure  $\Delta m_{23}^2$  and  $\sin^2 2\theta_{23}$  **PRL 97, 191801 (2006).**
  - High statistics constraints on alternative disappearance models.
     (e.g. neutrino decay, neutrino decoherence, sterile neutrinos ... ).

#### • Search for sub-dominant $v_e$ appearance.

– First observation or improved limit for small mixing angle  $\theta_{13}$ .

#### Atmospheric neutrino oscillations.

- Contained vertex  $v_{\mu}$  CC interactions. **PRD 73, 072002 (2006).**
- Neutrino-induced upward-going muons. PRD 75, 092003 (2007).

#### Cosmic ray physics.

– Muon charge ratio at TeV energies. arXiv/0705.3815 [hep-ex].

#### This Talk:

- New preliminary results on  $v_{\mu}$  disappearance based on exposure of
  - 2.5 x 10<sup>20</sup> protons on the NuMI target. arXiv/0708.1495 [hep-ex].
- New preliminary atmospheric muon and electron neutrino results.



#### **THE MINOS COLLABORATION**

Argonne • Athens • Benedictine • Brookhaven • Caltech • Cambridge • Campinas • College de France Fermilab • Harvard • IIT • Indiana • Minnesota Duluth • Minnesota Twin Cities • Oxford • Pittsburgh Rutherford • Sao Paulo • South Carolina • Stanford • Sussex • Texas A&M • Texas Austin Tufts • UCL • William & Mary • Wisconsin

#### **The MINOS Experiment**

- Accelerator beam of muon neutrinos produced by NuMI facility at Fermilab.
- Near Detector at Fermilab to measure spectrum and composition of beam.
- Far Detector at Soudan mine to study neutrino disappearance in beam.

735 km

Fermi Laboratory, Chicago

Soudan Mine, Minnesota

# The NuMI Beam





#### Neutrinos from the Main Injector (NuMI)

- 120 GeV protons from Main Injector directed onto 50g graphite target.
- 10µs spills with 2.4s cycle time.
- 2.5 x 10<sup>13</sup> protons per pulse.
- Typical beam power ~175 kW.
- Relative target position is moveable, making beam spectrum configurable.
- Majority of running in LE configuration.



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### **The MINOS Detectors**







#### **Near Detector**

1 kT mass 1 km from target 282 steel planes 153 scintillator planes 100m underground

#### **Functionally Identical Detectors**

steel and scintillator sampling calorimeters.

Magnetized steel (B ~1.3T).

GPS time-stamping for synchronization.

#### **Far Detector**

5.4 kT mass 735 km from target 486 steel planes 484 scintillator planes 700m underground

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**Look for** 
$$v_{\mu}$$
 **deficit**:  $P(v_{\mu} \rightarrow v_{\mu}) = 1 - \sin^2 2\theta \sin^2 \left(\frac{\Delta m^2 L}{E}\right)$ 





# 2+ Years of NuMI Running







#### Improvements over 2006 analysis:

- Better reconstruction.
- Improved event selection.
- Improved shower modelling.
- New intra-nuclear modelling.
- CC/NC interactions separated using multivariate 2D likelihood procedure combining information from:
  - Track observables.
  - Event length.
  - Event kinematics.

Improvement in selection ~1% more CC signal. ~50% less NC background.

Data and Monte Carlo agree well.



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- Parameterize Fluka 2005 hadron production model as  $f(x_F, p_T)$ .
- Fit to near detector data collected in different beam configurations.
  - incorporate into the fit: horn focusing current, beam misalignments, cross-sections, neutrino energy scale, neutral current background.
- Improved agreement between data and MC in all configurations.





# **Predicting the Far Spectrum**



#### • Directly use near detector data to extrapolate from near to far detector.

- Use Monte Carlo to correct for energy smearing and detector acceptance.
- Use a beam transfer matrix derived from the beam simulation to relate neutrino interactions in each detector via their parent hadrons.





## **Systematic Uncertainties**



- Systematic uncertainties on oscillation parameters evaluated by fitting fake data sets generated from MC with systematic shifts applied.
- The three largest uncertainties identified from this study are included as nuisance parameters in the oscillation analysis.

| Uncertainty                    | Δm² (10 <sup>-3</sup> eV²) | sin² 2θ |
|--------------------------------|----------------------------|---------|
| Near/far normalization (4%)    | 0.065                      | <0.005  |
| Abs. shower energy scale (10%) | 0.075                      | <0.005  |
| NC normalization (50%)         | 0.010                      | 0.008   |
| All other systematics          | 0.040                      | <0.005  |
| Total uncertainty (quad. sum)  | 0.11                       | 0.008   |
| Statistical uncertainty        | 0.17                       | 0.080   |





#### PRELIMINARY OSCILLATION RESULTS FOR 2.5x10<sup>20</sup> POTs DATA.



| Data sample              | Observed | Expected<br>(no osc.) | Observed /<br>Expected |
|--------------------------|----------|-----------------------|------------------------|
| $ u_{\mu}$ (all E)       | 563      | 738 ± 30              | 0.74 (4.4σ)            |
| ν <sub>μ</sub> (<10 GeV) | 310      | 496 ± 20              | 0.62 (6.2σ)            |
| ν <sub>μ</sub> (<5 GeV)  | 198      | 350 ± 14              | 0.57 (6.5σ)            |

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## **Allowed Parameter Space**



#### MINOS Preliminary



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## **Atmospheric Neutrinos at MINOS**



# • MINOS far detector can be used to study atmospheric neutrinos.

- 5.4 kT mass generates high rate of atmospheric neutrino interactions.
- 700m depth provides shielding against cosmic muon background.
- magnetic field enables separation of neutrino and anti-neutrino events.
- Calorimeter detector design enables measurement of total energy.
- Atmospheric neutrino analyses:
  - contained vertex interactions.PRD 73, 072002 (2006).
  - neutrino-induced up-going muons.
     PRD 75, 092003 (2007).
  - Reported here are new preliminary contained vertex muon and electron atmospheric neutrino results.



Contained vertex events classified as follows:

- Fully Contained (FC).
- Down-Going Partially Contained (PCDN).
- Up-Going Partially Contained (PCUP).



## **Atmospheric Muon Neutrinos**



- Updated contained vertex  $v_{\mu}$  analysis based on exposure of 12.23 kT-Yrs.
- Observe 277 events, with expectation of 354 ± 47 in absence of oscillations.
- Select events with well-measured muon direction based on timing information.

- 105 downward-going, 77 upward-going.

 $R_{up/down}^{data} / R_{up/down}^{MC} = 0.72_{-0.11}^{+0.13} (stat) \pm 0.04 (sys)$ 

• Select events with well-measured muon charge based on curvature in B-field.

– 112 neutrinos, 55 anti-neutrinos.

 $R_{\overline{\nu}/\nu}^{\text{data}} / R_{\overline{\nu}/\nu}^{\text{MC}} = 0.93^{+0.19}_{-0.15}(\text{stat}) \pm 0.12(\text{sys})$ 

 Oscillation analysis carried out by binning events according to their Bayesian *L/E* resolution.



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- Preliminary MINOS atmospheric v<sub>e</sub> results based on exposure of 6.18 kT-Yrs.
  - Identify  $\nu_{\rm e}$  events as contained vertex electromagnetic showers.
  - Observe 89 candidate  $\nu_{\rm e}$  events with expectation of 89  $\pm$  17 events.
  - Observe 113 candidate  $\nu_{\mu}$  events with expectation of 150  $\pm$  30 events.

 $R_{\mu/e}^{data} / R_{\mu/e}^{MC} = 0.74_{-0.10}^{+0.12} (stat) \pm 0.05 (sys)$ 

- Use selected  $v_e$  event sample to measure atmospheric neutrino flux normalization relative to *Bartol04* flux model.
  - account for oscillations of true  $v_{\mu}$  events in selected  $v_{e}$  event sample.

 $S_{atm} = 1.07 \pm 0.12(stat) \pm 0.08(sys)$ 











- MINOS has had a successful second year of beam running.
  - 3.6x10<sup>20</sup> PoTs have now been accumulated after two years.
- Updated oscillation measurement based on 2.5x10<sup>20</sup> PoTs.

$$\left|\Delta m_{32}^2\right| = 2.38^{+0.20}_{-0.16} \times 10^{-3} \text{ eV}^2$$
  
 $\sin^2 2\theta_{23} = 1.00_{-0.08}$ 

Other oscillation analyses using beam data are progressing.

 $-v_e$  appearance, anti- $v_{\mu}$  disappearance, sterile neutrinos...

- Updated atmospheric muon and electron neutrino results.
  - developing combined analysis of all MINOS atmospheric neutrino data.

See poster session for more information on latest MINOS results

# **Backup Slides**

NuMI Beam Line

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# **Main Injector**

22.420

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Fermi Laboratory









#### **PMT Dark Box**



#### **Scintillator Strips**

### WLS Fibres









# **MINOS Calibration**

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– PMT gain and linearity.

#### Cosmic ray muons:

- relative strip calibration.
- intra-detector calibration.

| Calibration Error:   |      |  |
|----------------------|------|--|
| - ND calibration:    | 3.1% |  |
| - FD calibration:    | 2.3% |  |
| - ND/FD calibration: | 3.8% |  |

#### Overall Energy Scale:

- Calibration detector at CERN measured  $e/\mu/\pi/p$  response.

#### Energy Resolution (E in GeV):

- Hadrons: 56%/ $\sqrt{E} \oplus 2\%$
- Electrons: 21%/ $\sqrt{E} \oplus 4$ %/E



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### **Reconstruction of a MINOS Event**



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# **Event Topologies in MINOS**



v<sub>µ</sub> CC Event μ-W p n υz ٧Z 3.5m վկակեսես

long µ track & hadronic activity at vertex

 $\frac{\text{NC Event}}{v_{\alpha}}$ 



υz





short event, often diffuse

v<sub>e</sub> CC Event





# short, with typical EM shower profile

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- CC events selected using a likelihood based procedure with six input PDFs that show discrimminating power between true CC and NC interactions:
  - Number of track planes.

  - Track pulse height per plane. Reconstructed y  $(E_{shw}/E_{y})$ .
- Goodness of muon track fit.
  - Number of track only planes.
     Reconstructed muon charge.
- 2D PDFs are used to take account of correlations with event length.
- The discrimminant variable (PID) is defined as follows:

$$P_{CC}(X,Y,Z,...) = P(X|CC) P(Y|CC) P(Z|CC) ... P(CC)$$

 $P_{NC}(X,Y,Z,...) = P(X|NC) P(Y|NC) P(Z|NC) \dots P(NC)$ 

$$PID = \frac{P_{CC}}{P_{CC} + P_{NC}}$$

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Good agreement between data and Monte Carlo observed for these variables.



### **Near Detector Interactions**



- High event rate in near detector. – Multiple interactions per spill.
- Events separated based on topology and timing.
  - Timing resolution ~20 ns
  - Spatial resolution ~4 cm
- No significant bias in event rate.





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### **Far Detector Interactions**



- Beam interactions identifiable with "spill trigger".
  - GPS spill time is sent via internet from near to far detector.
  - Events within  $\pm 50 \mu s$  of spill written out by far detector DAQ.





### **Far Detector Timing**



#### **MINOS PRELIMINARY**





### **Far Detector Event Rates**







## **Far Detector Distributions**



Far detector data is well described by oscillation best fit



### **Far Detector Distributions**





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### Far Detector Backgrounds



#### **MINOS PRELIMINARY**





## **Comparison with 2006 Result**

#### (PRL 97, 191801)



**MINOS Preliminary** 



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The MINOS Experiment, slide 37



### **Comparison with 2006 Result**



#### (PRL 97, 191801)



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### Changes from 2006 Result (PRL 97, 191801)







## **Comparison of Runs I and IIa**





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## $\underline{v}_{\mu}$ **Disappearance**



#### **MINOS Sensitivity as a function of Integrated POT**





# NC Analysis



- Neutral current interactions are unaffected by standard oscillations, so can be used to constrain oscillations into sterile neutrinos.
- Define sterile mixing parameter  $f_s$  as the fraction of disappearing muon neutrinos that oscillate into sterile neutrinos.



• Far detector data for this analysis currently blinded – analysis in progress.



### <u>v</u>e Appearance

δ (π)



- MINOS can constrain or measure  $\theta_{13}$  by searching for  $v_e$  appearance.
- Challenges are to separate signal and understand background.
  - NC events form dominant background.
  - Much effort has gone into developing techniques for distinguishing between electromagnetic and hadronic showers.
  - Data-driven techniques for background determination also in development.
- MINOS sensitivity will soon be comparable with the current world best limit (CHOOZ).

90% CL Sensitivity to sin<sup>2</sup>(2013) 2 MINOS  $\Delta m_{23}^2 = 2.7 \ 10^{-3} \ eV^2$ 1.8  $sin^{2}(2\theta_{23}) = 1$ 1.6 4x10<sup>20</sup> pot To be superseded soon 1.4 CHOOZ 1.2 90% CL Excluded 1 0.8 0.6 0.4  $\Delta m^2 < 0$ 0.2 0 -2 -1 10 10

sin<sup>2</sup>(2013)